



WISH

The Cosmic Dawn Treader

Wide-field Imaging Surveyor for High-Redshift

超広視野初期宇宙探査衛星

WISH Working Group

<http://www.wishmission.org/en/index.html>

1. WISH Summary
2. Search for Very High- z Galaxies
3. SNe
4. Other Science
5. WISH System

WISH WG Members

Toru Yamada, K.Mawatari, M.Kubo (Tohoku University)

Ikuru Iwata, S.Tsuneta,

T.Kodama, Y.Komiyama (NAOJ)

H.Matsuhara, T.Wada, Y.Oyabu (JAXA/ISAS)

H.Sugita, Y.Sato, A.Okamoto (JAXA)

K.Ohta, K.Yabe , Tsutsui (Kyoto University)

T.Morokuma, Chihiro Tokoku

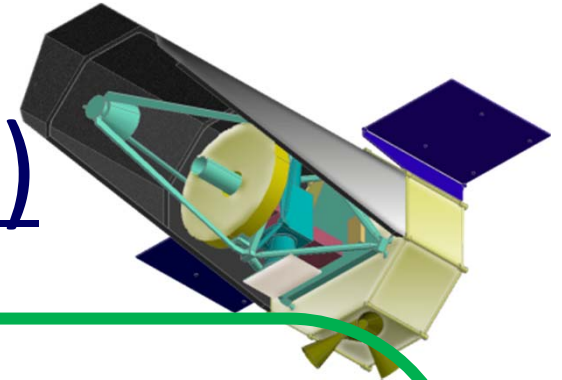
M.Doii, N.Yasuda (University of Tokyo)

N.Kawai (TiTEC), Yonetoku (Kanazawa U),

A.Inoue (Osaka Sangyo U.)

Y.Ikeda (Photocoding), S.Iwamura (M.R.J)

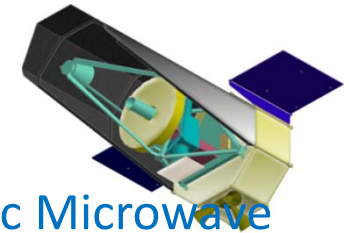
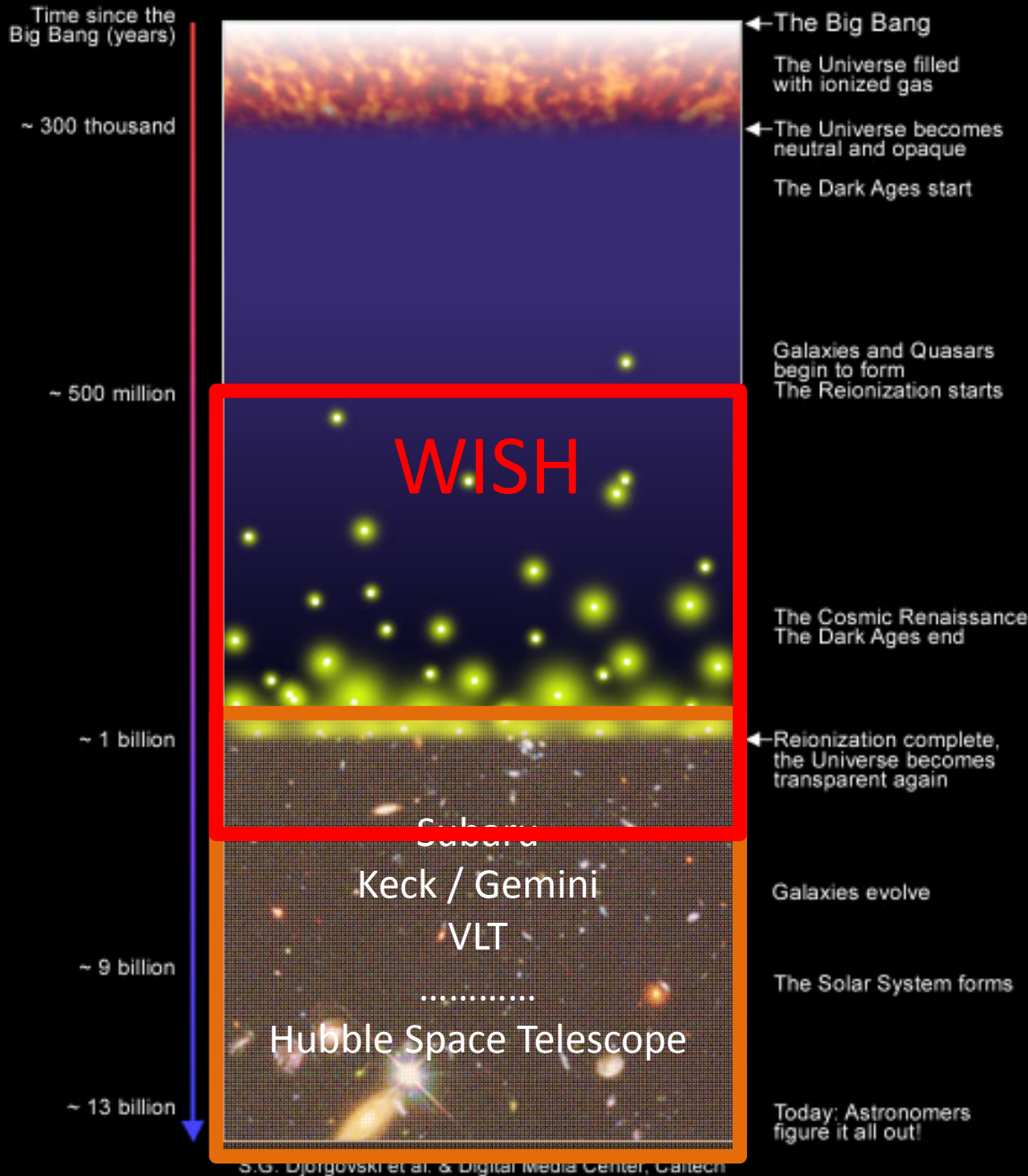
WISH Brief Summary (1)



- NIR Deep and Wide-field Imaging Surveyor
- Exploring the 1st generation galaxies
- Dedicated, $\sim 100 \text{ deg}^2$, 28AB ($\sim 25 \text{ nJy}$)
- Concept developed under JAXA/ISAS WG
(the WG was selected in Sept 2008)
to be launched in late 2010's (NET2018)
- Utilizing heritage of existing technology

What is the Reionization Era?

A Schematic Outline of the Cosmic History



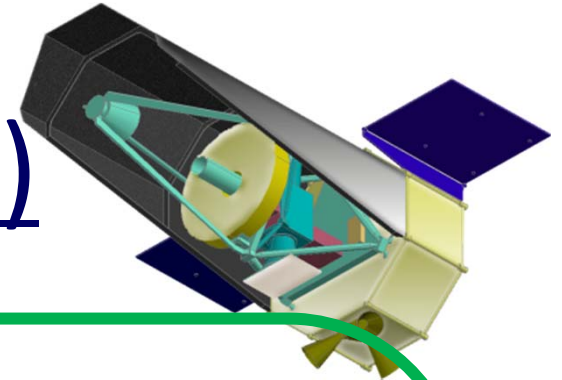
Cosmic Microwave Background (CMB)
Universe: **Neutral**

First-Generation Galaxies

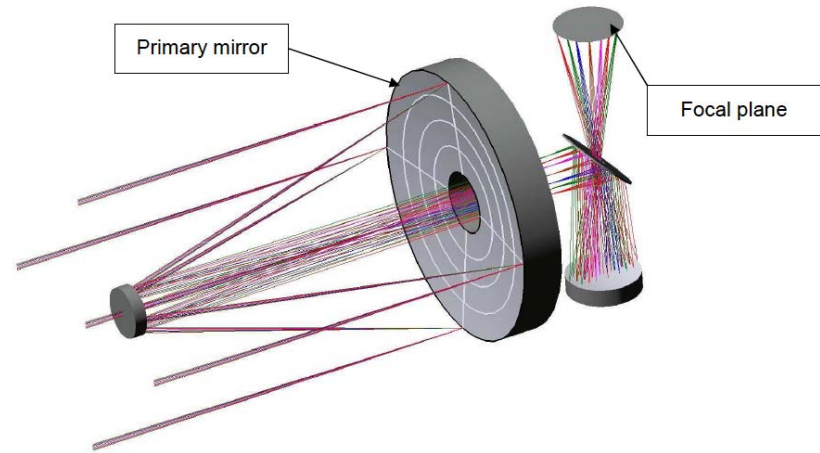
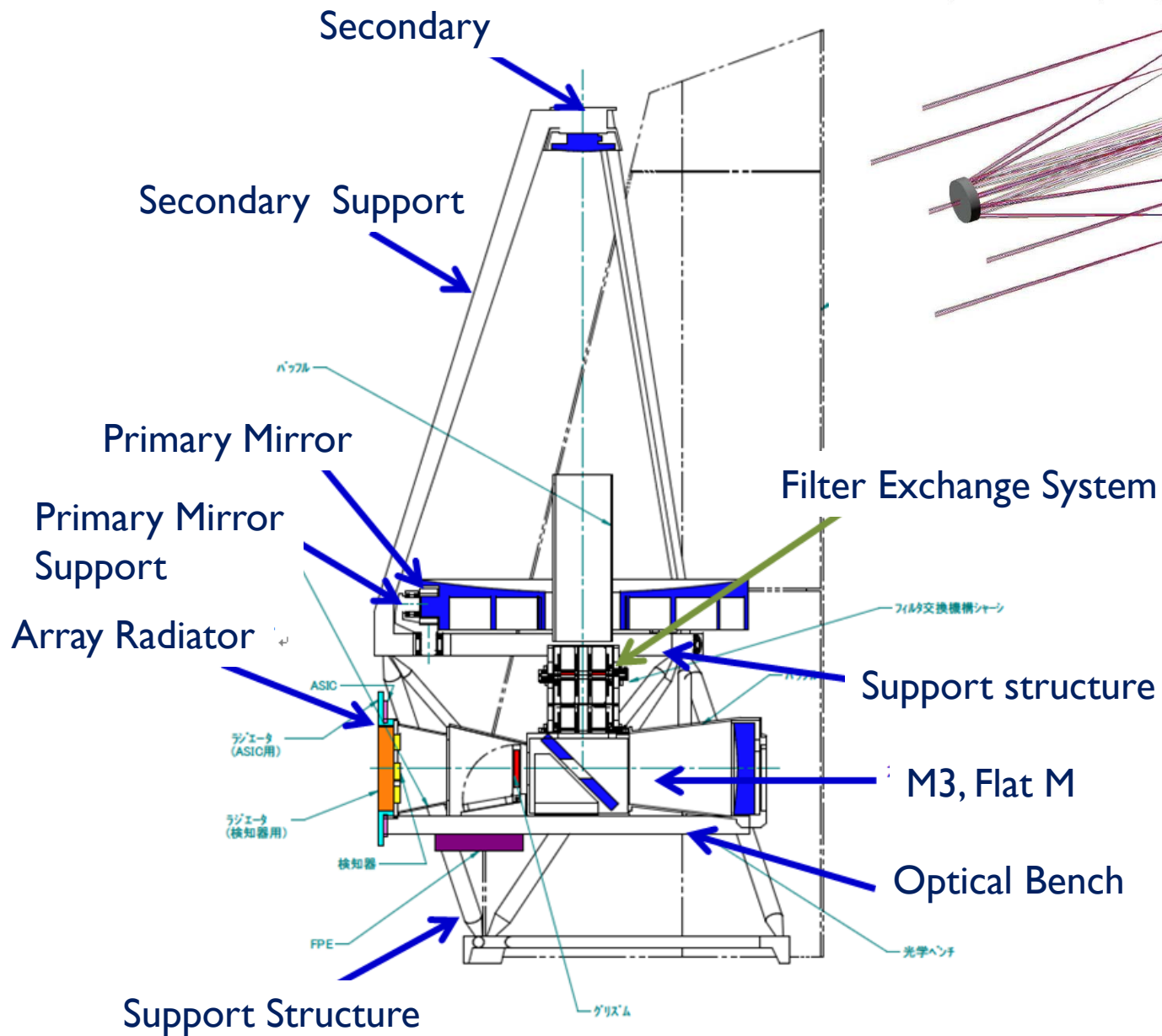
Ultimate Frontier of Galaxies

Universe: **Ionized**

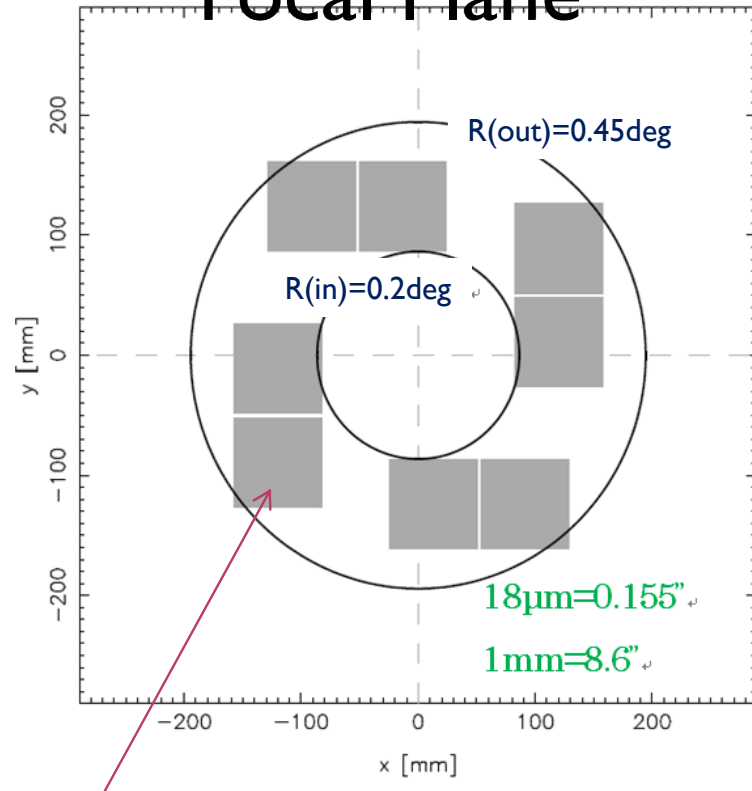
WISH Brief Summary (2)



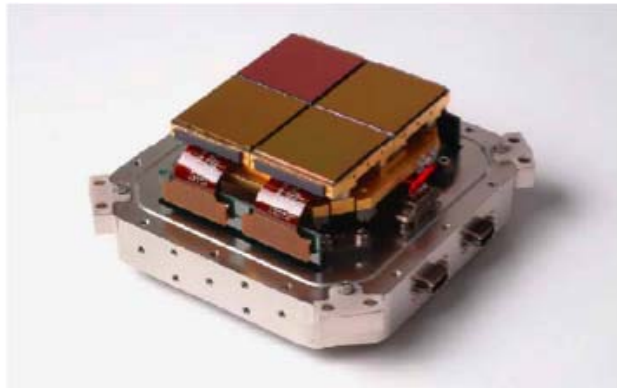
- 1-5 μm wavelength range
- 1.5m diameter telescope
- Very Wide-Field Imager
 - ~900 arcmin^2 FoV
- pixel scale: 0.155" / 18 μm (f/16)
- Cooled to 90-100K (telescope)
- SE-L2, JAXA HIIA



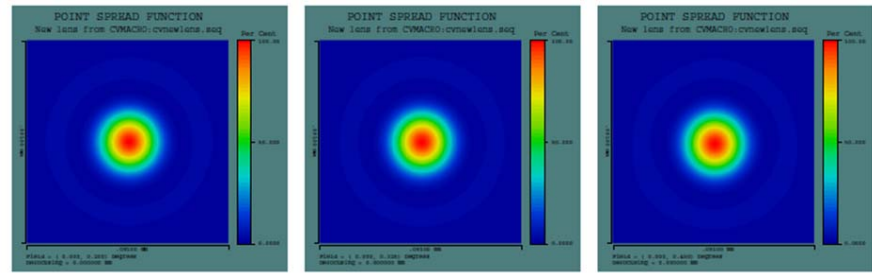
Focal Plane



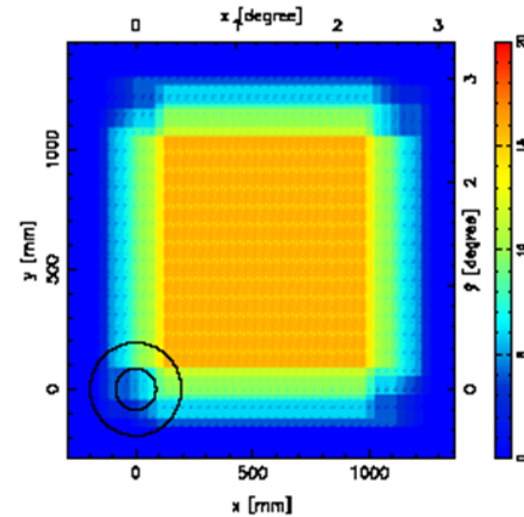
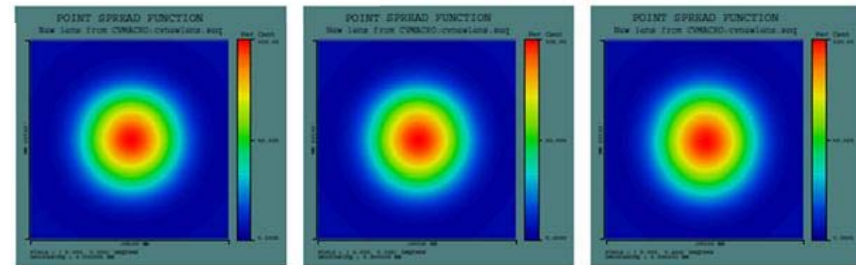
4 x 2kx2k FPA



$R=0.2, 0.325, 0.4 \text{ deg @ } 1.25 \mu\text{m}$



$R=0.2, 0.325, 0.4 \text{ @ } 2.2 \mu\text{m}$

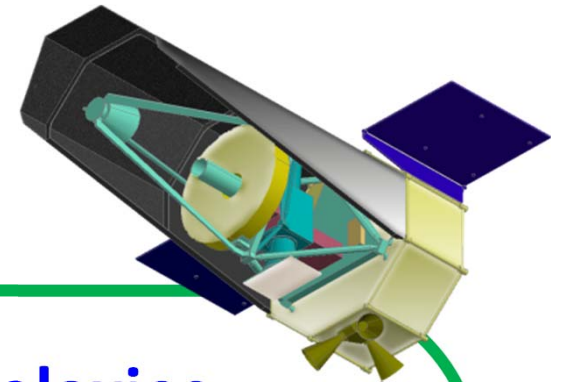


WISH: Survey Strategy

Survey categories

	Depth (3σ) (AB mag)	Area	Example of the Filters (a plan, to be determined)
Ultra Deep Survey (UDS)	28	100 deg ²	1.4,1.8, 2.3, 3.0 μm
Multi-Band Survey (MDS)	28	10 deg ²	1.0,4.0
Ultra Wide Survey (UWS)	24-25	1000 deg ²	1.4, 1.8, 2.3
Extreme Survey	29-30	0.25 deg ²	1.0, 1.4, 1.8

WISH Science Goals (1)



- **Exploring the Ultimate Frontier of Galaxies**
Detection of 1st-generation galaxies and studying cosmic reionization over $z=7-15$
- **NIR search and light curves for type-Ia SNe**
History of cosmic expansion and Dark Energy
- Transients: high- z GRB, luminous SNe
- Huge statistics and New discovery

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WISH: Survey Strategy

Surveys achieved within ~ 1000 days (50% overhead)

WISH can detect

$\sim 10^4$ galaxies at $z=8-9$,

$\sim 10^{3-4}$ galaxies at $z=11-12$,

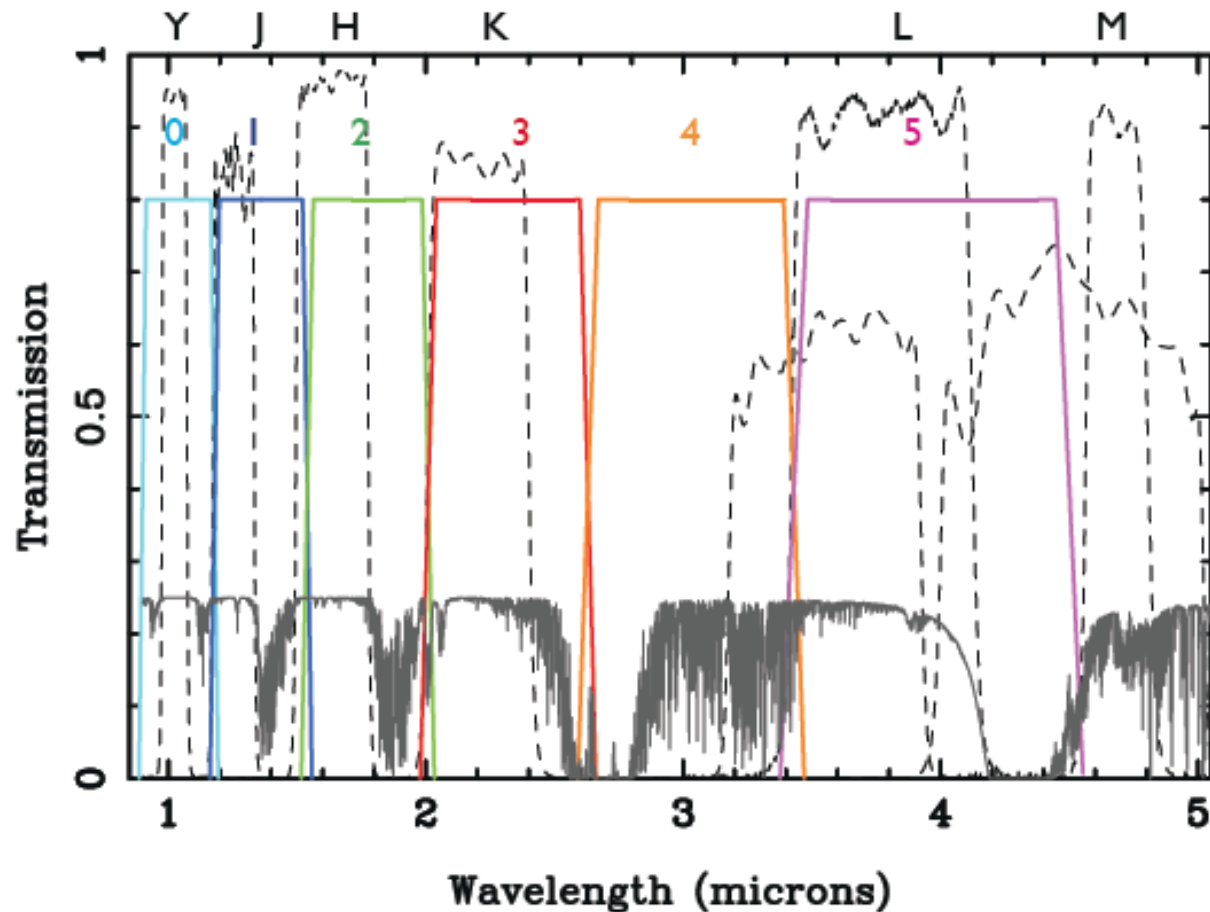
and

$\sim 50-100$ galaxies at $z=14-17$

Many of them are feasible spectroscopic targets

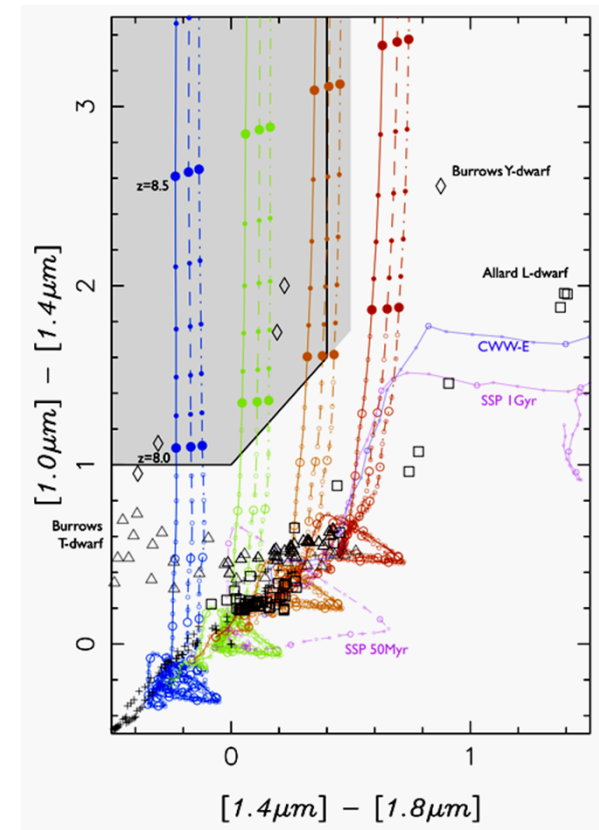
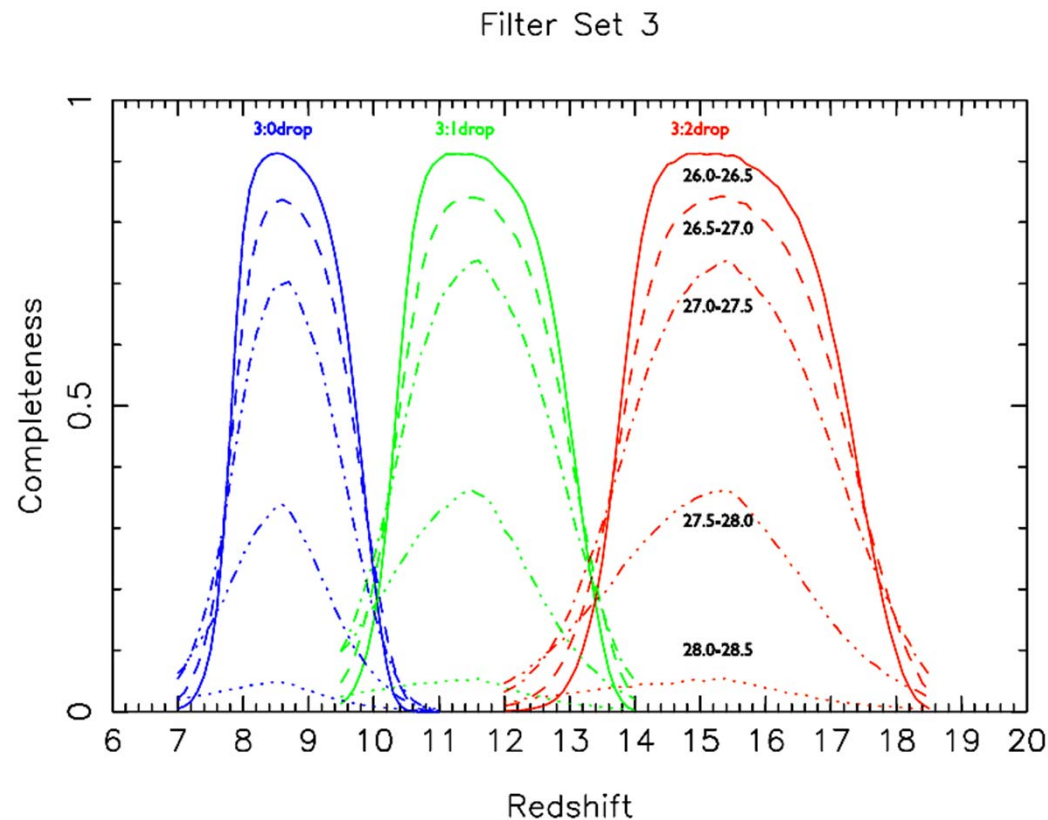
WISH: Survey Strategy

set of broad-band filters to select high-z galaxies



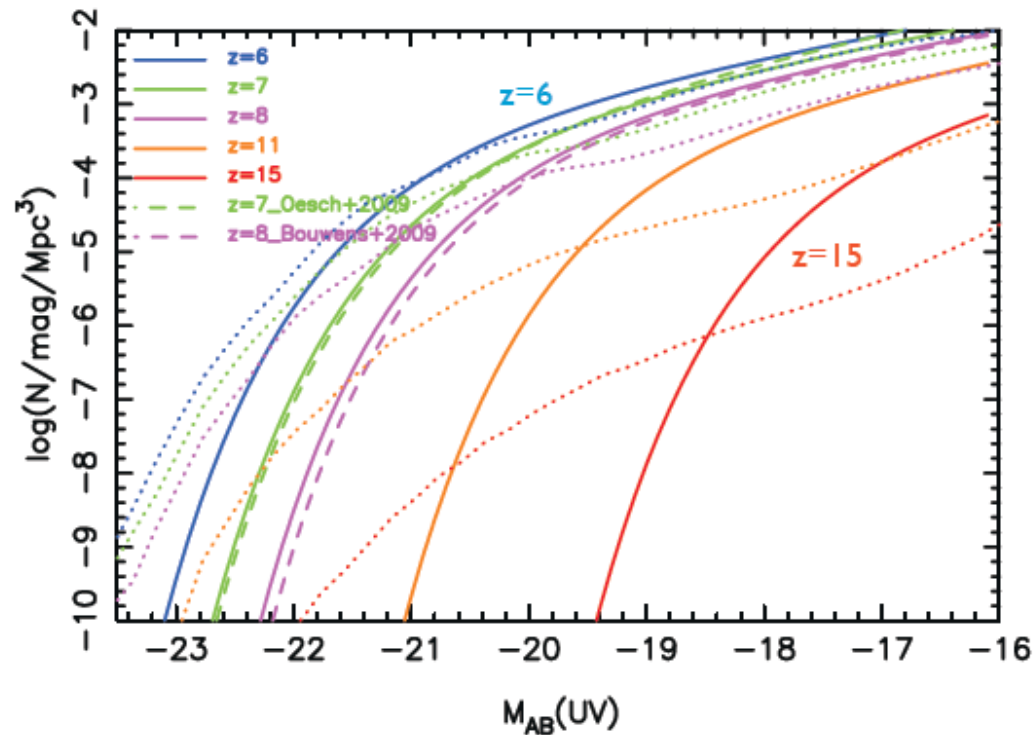
WISH: Survey Strategy

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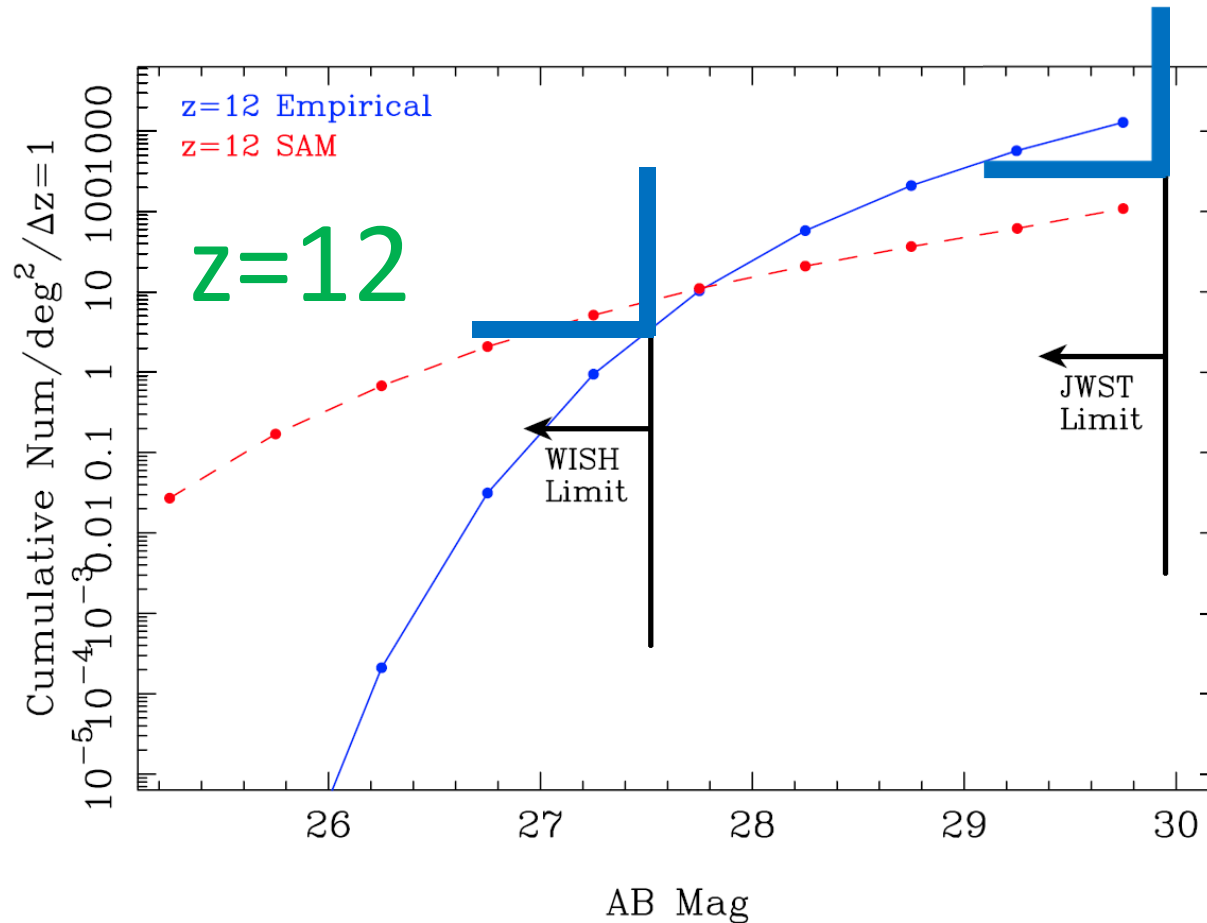
How deep, how large area should we observe?

Observed (z=6-8, HST WFC3 **dashed lines**) and
Predicted (z=6-15) UV Luminosity Function of Galaxies



Solid lines: empirical expectation (extrapolation) from z=6-8 luminosity function
Dotted lines: expectation based on galaxy models (semi-analytic treatment)

In the observers' frame...



Limit of
AB>27-28 is
needed

—
Empirical
luminosity evolution

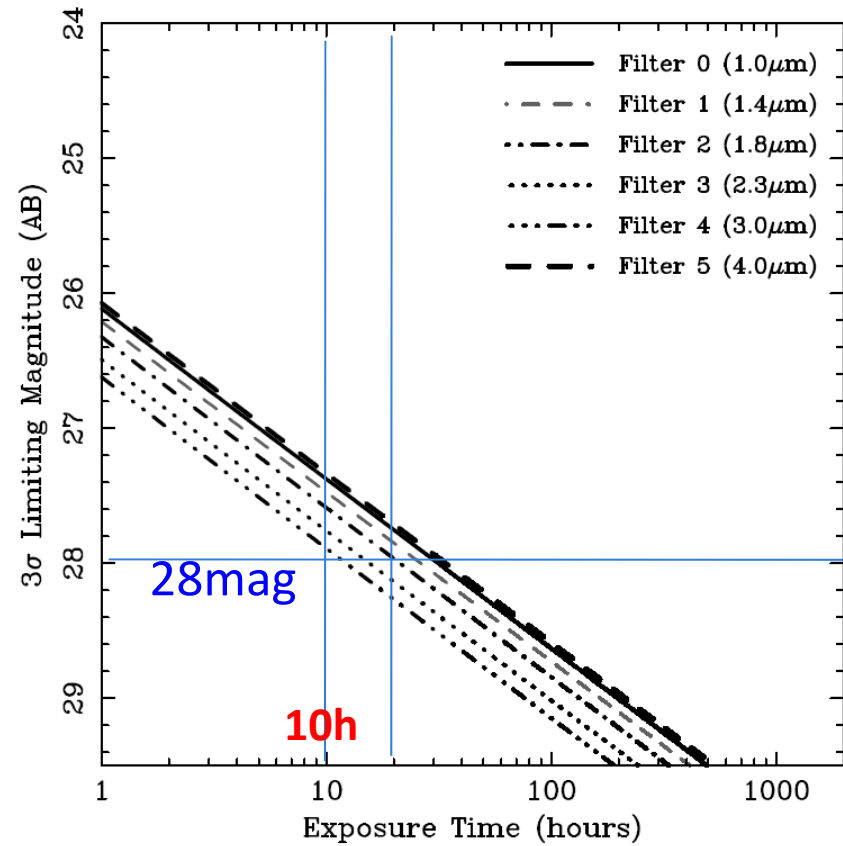
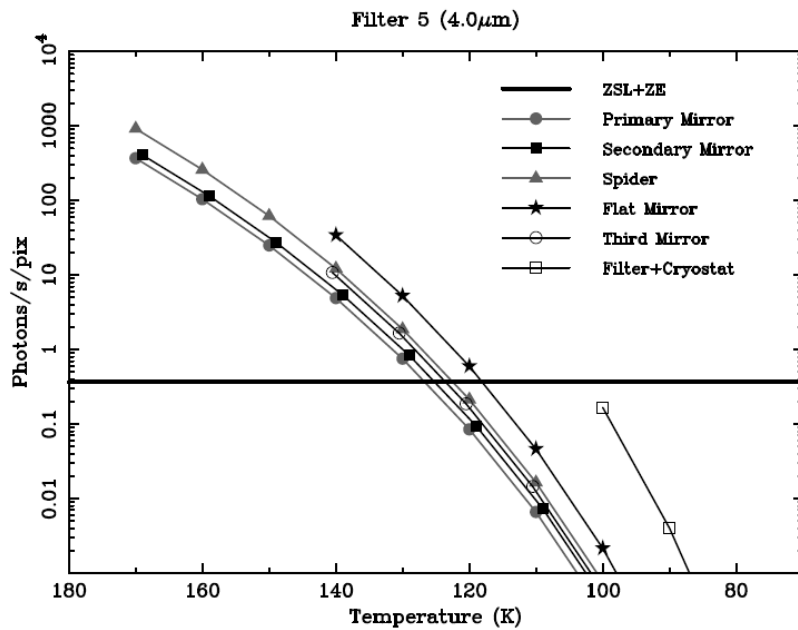
- - -
Semi-analytic
Model prediction

FoV JWST NIRCam 2.2'x2.2' x 2ch (per filter)

~ 2.8x10⁻³ deg²

WISH: Survey Strategy

Telescope should be cooled to $\sim 100\text{K}$
(detector to $\sim 40\text{K}$ for $5\ \mu\text{m}$)



WISH Science Goals

Expected number of the *observed* very high-redshift galaxies

		Number Density [objects per 1 deg ²] for AB < 28.0			
	redshift	No Evolution	Empirical	SAM	DMH
1.0μm-drop	8-9	4,000	1,700	630	850
1.4μm-drop	11-12	2,400	100	50	4.1
1.8μm-drop	14-17	1,200	0.72	1.1	0.003

Numbers for 1 deg² , <28AB

Galaxies bright enough for deep spectroscopy with ELT + AO spectrograph

WISH: Survey Strategy

Surveys achieved within ~ 1000 days (50% overhead)

WISH can detect

$\sim 10^4$ galaxies at $z=8-9$,

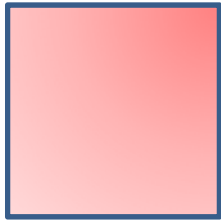
$\sim 10^{3-4}$ galaxies at $z=11-12$,

and

$\sim 50-100$ galaxies at $z=14-17$

Many of them are feasible spectroscopic targets

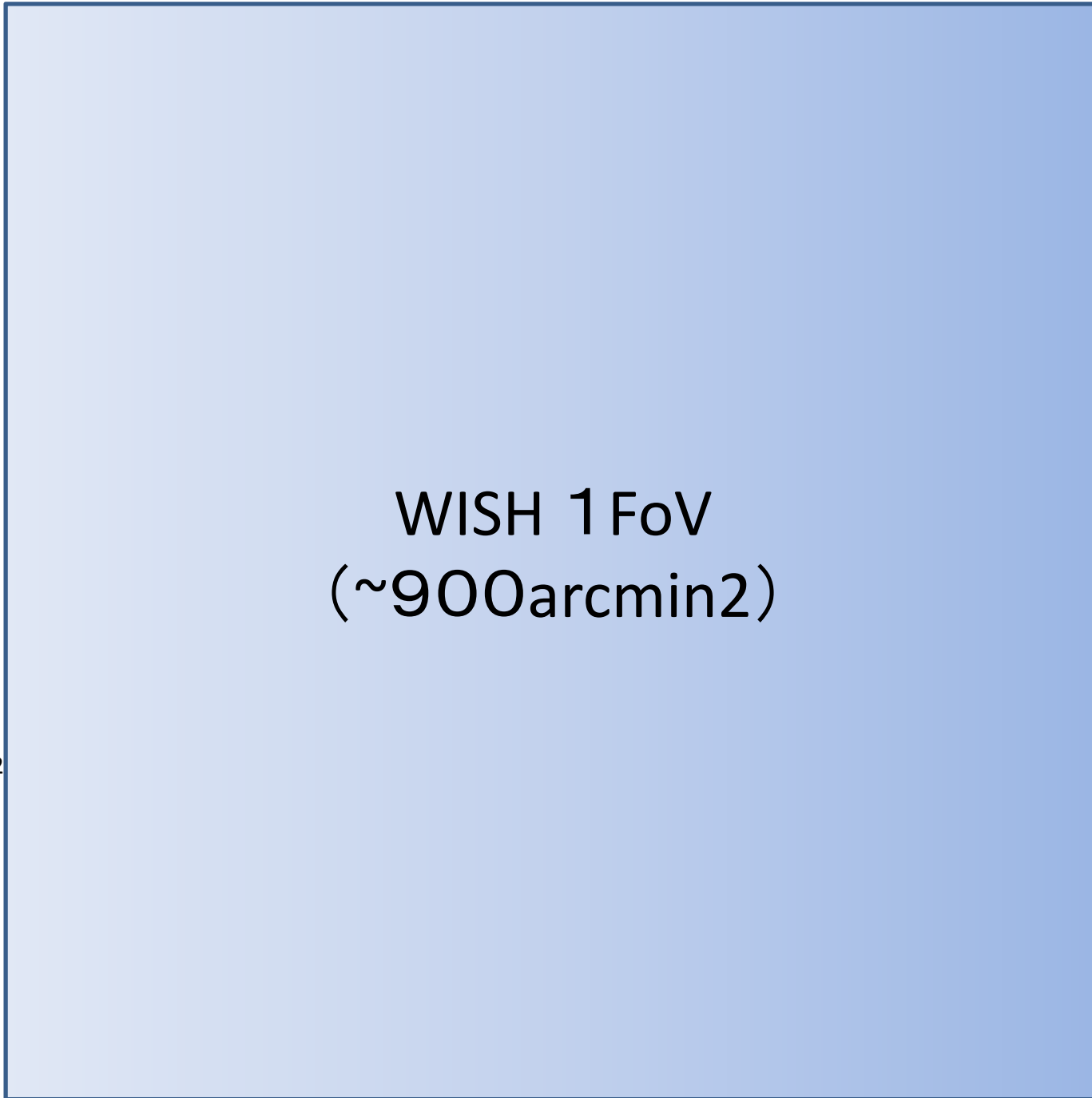
HST UDF



~ JWST
(2x2.2x4.3
=19 arcmin²)

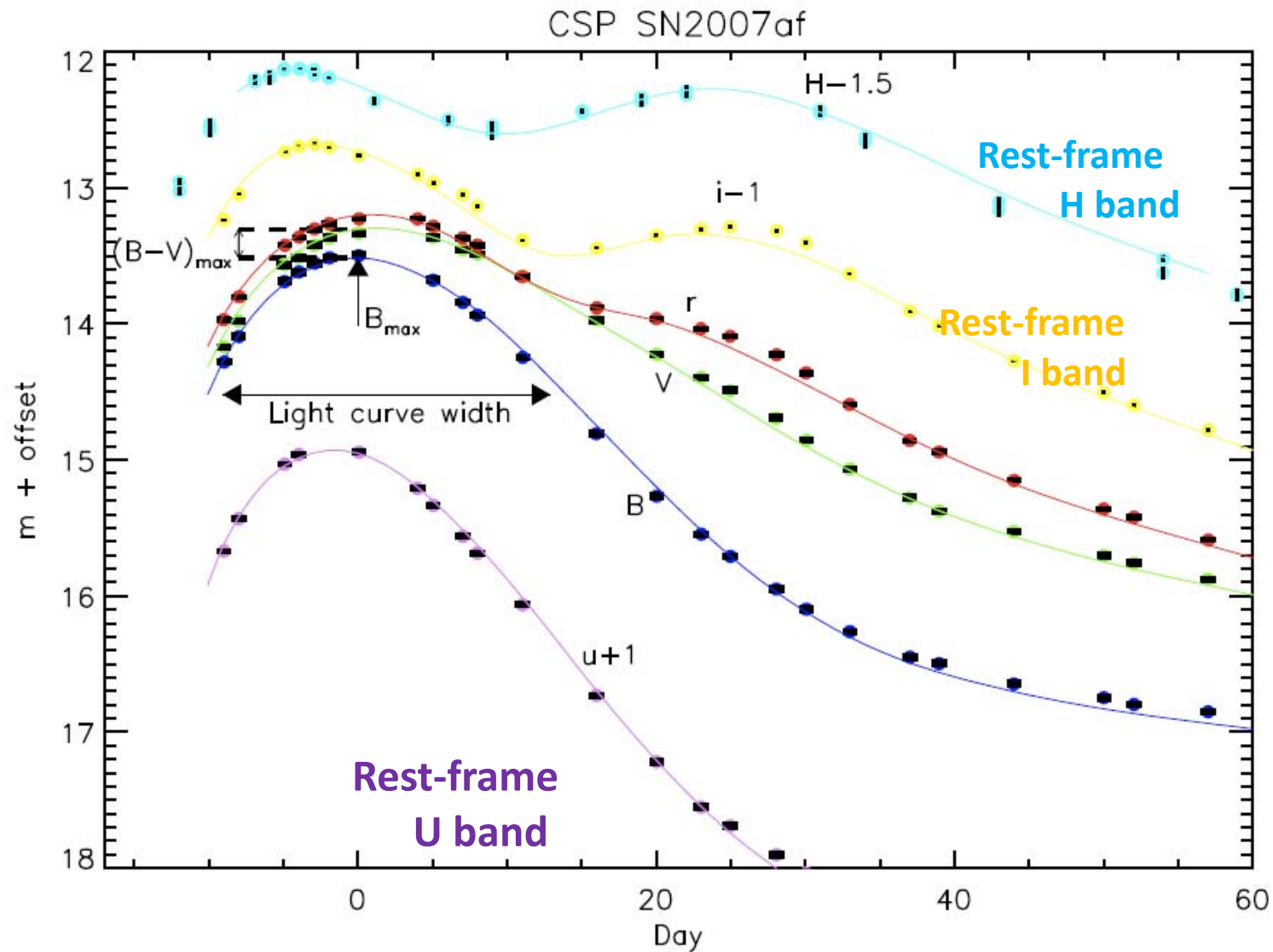
~SPICA
FPC-S
5'x5'=25arcmin²

WISH 1 FoV
(~900arcmin²)

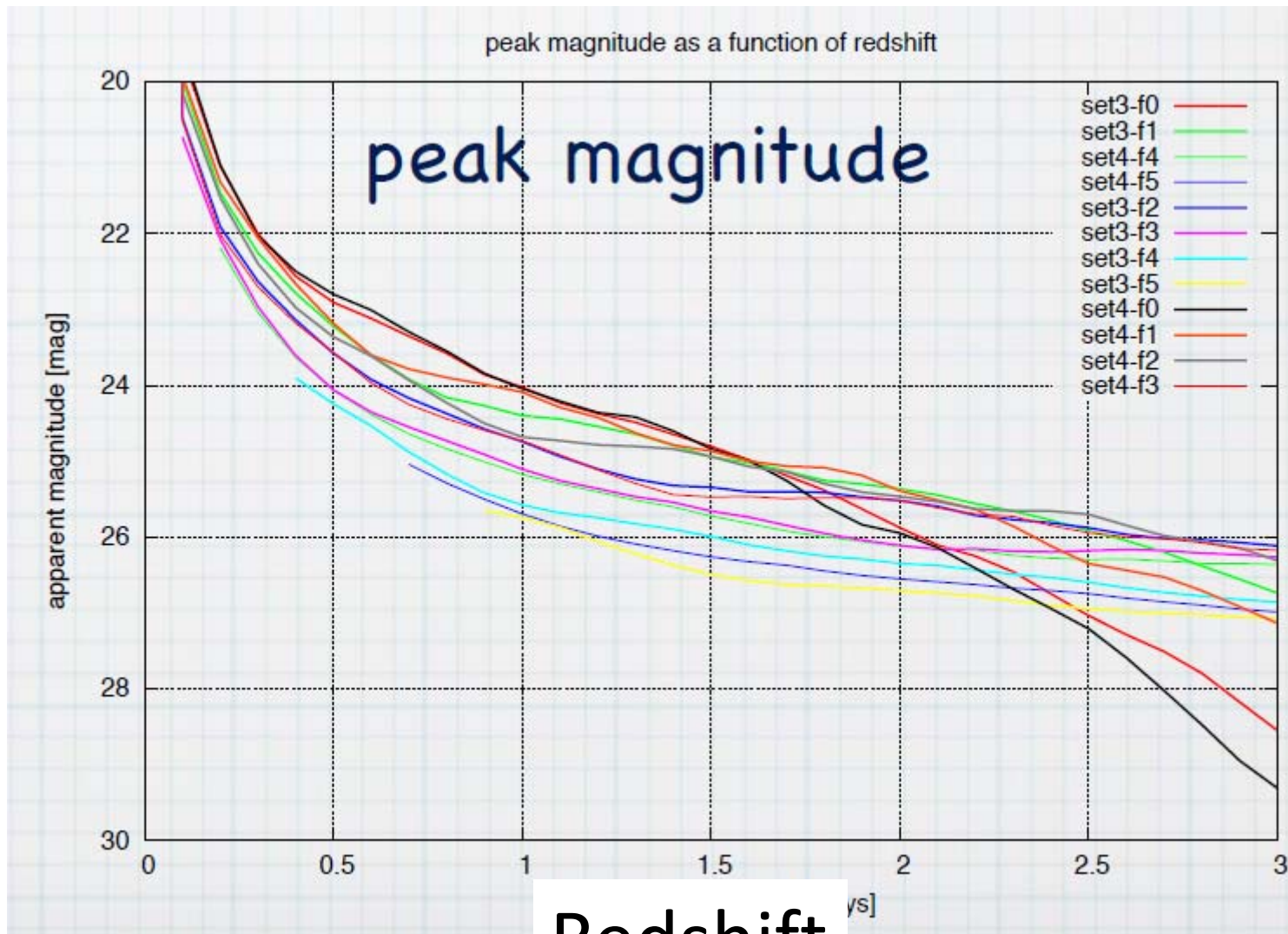


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Light Curves of Type Ia SNe



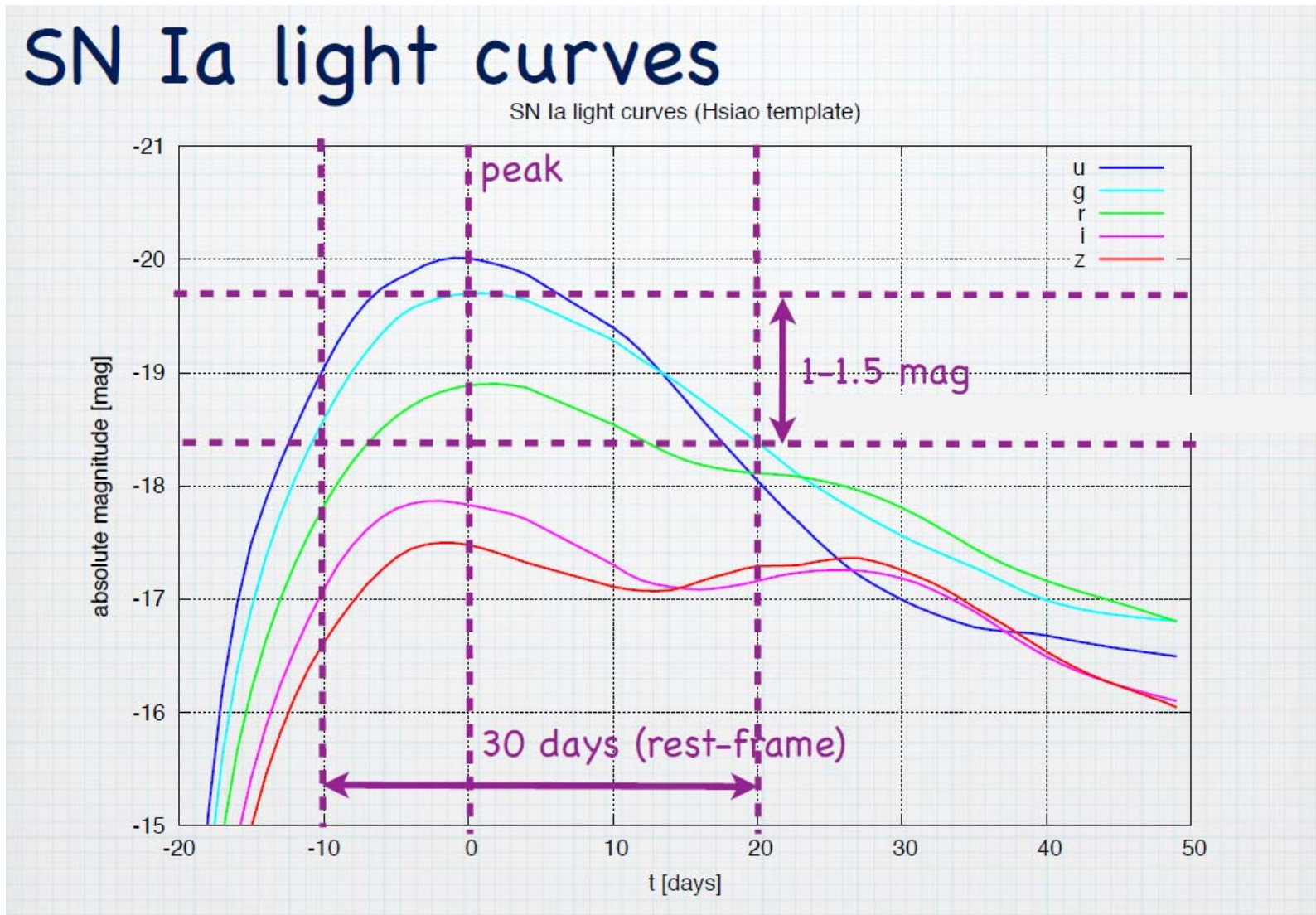
From Goobar and Leibundgut 2011



Redshift

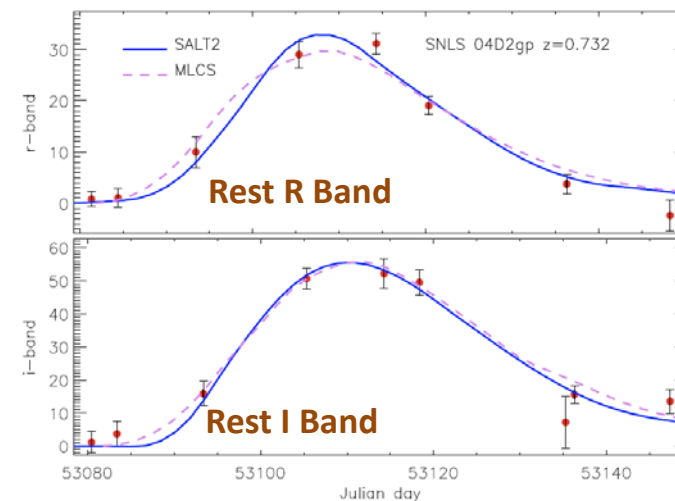
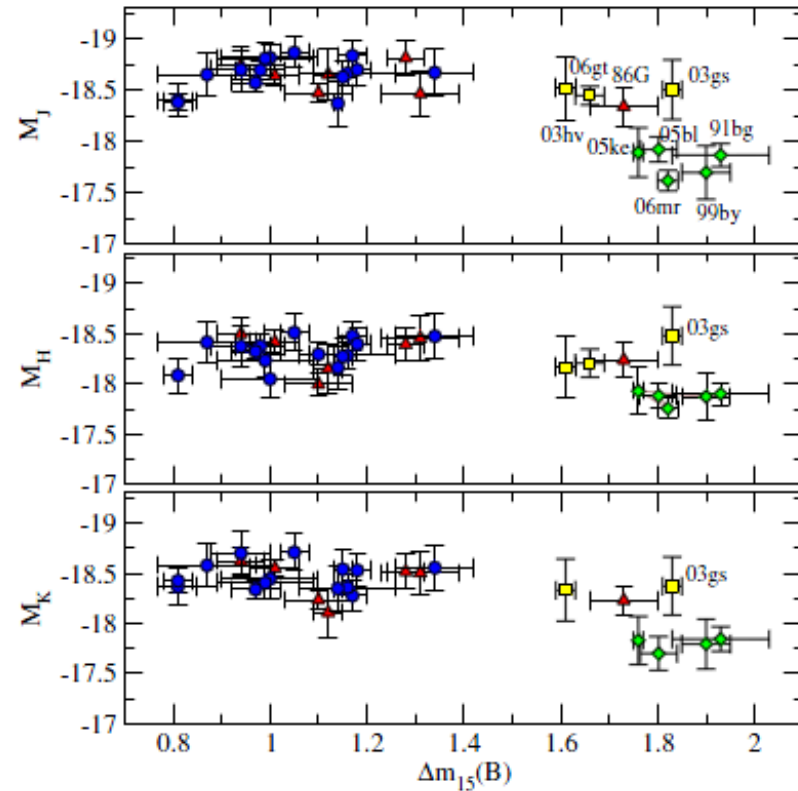
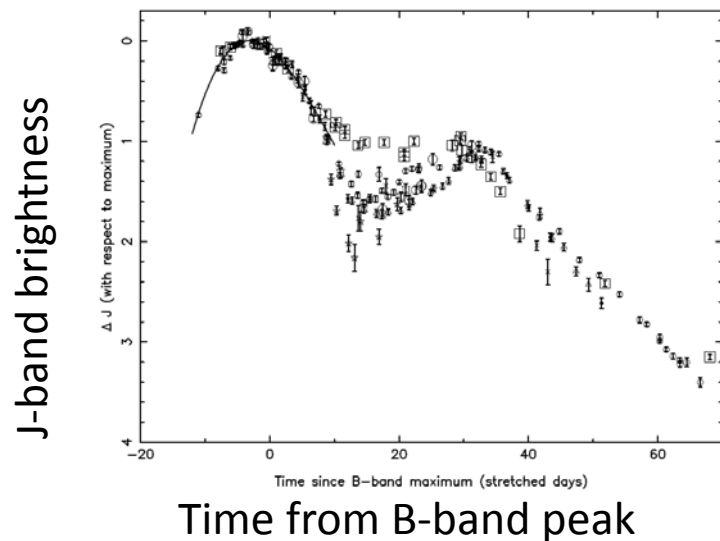
Morokuma, 2010

WISH type-Ia SNe



Advantage: rest-frame IR

- Small scatter in peak magnitude
n.b. more dispersion for $T_{\text{peak}}(\text{IR}) > T_{\text{peak}}(\text{B})$
- Double peaks in Light Curve.
First peak can be fitted by a single template
- Less sensitive to the Fitting Models of LC



From Goobar and Leibundgut 2011

WISH type-Ia SNe

Ultra Deep Survey 3-4 band AB28mag

N=5-10 times/year (@ $z \sim 1$... $\Delta t \sim 10$ days)

→ magnitude limit 27.1 (N=5) – 26.8 (N=10)

1mag margin: $m < 26.1$ (N=5) $m < 25.8$ (N=10)

	Rest-frame I Band	Rest-frame H Band
N=5	$z=0.2-2.2$ 2000 SNIa in 80 deg ²	$z=0.-1.4$ 2000 SNIa in 170 deg ²
N=10	$z=0.2-1.6$ 2000 SNIa in 22 deg ²	$z=0.-1.0$ 2000 SNIa in 67 deg ²

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- Galaxy Formation and Evolution
- Transient
 - GRB toward $z=20$, type-IIIn SNe
- Variability Survey
 - AGN in High- z Galaxies
- Strong Lensing
 - Double Einstein Rings
- Quasar toward $z=20$ (Ultra Wide Survey)
- Weak Lensing Cosmology
 - (Ultra Wide Survey)

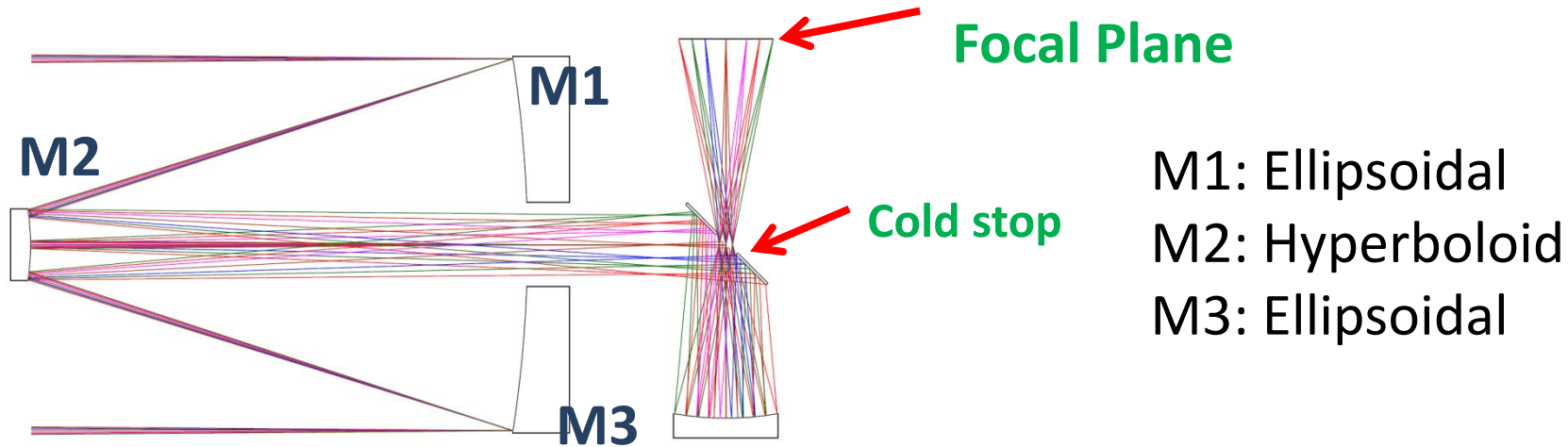
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Overall Parameters		
Launch Date (plan)	NET 2017	Mission Life Time > 5yr
Launcher	H-IIA 4/4D-LC lower case PAF 2360S	Capable to launch 1.5t (dual launch, SE-L2)
Size	φ3.3m x 5.2m	Launch Configuration Including the startracker
Mass	@launch (WET) 1.3t	margin 0.2t
Orbit	Sun Earth L2 Halo	
Power	Max 1.2 kW	SAP power 1.6kW (BOL) 1.4kW (EOL)

Mission Part		
Optics	M1 φ1.5m ellipsoidal M2 φ0.28m hyperboloid M3 φ0.41m ellipsoidal Deff..limit at 1-5μm Fov φ0.2-0.7deg 0."15/18μm(pixel)	<ul style="list-style-type: none"> • Cooled to 100K • flat mirror also works as the cold stop • light going through the flat mirror hole: option • slitless grism spectroscopy at 1-2.5μm
Focal Plane Arrays	HgCdTe (HAWAII-2RG) 2K x 2K 32 chips (128Mpix)	0.9-5.3μm 1pixel=18μm
Filter Exchanger	Flip-type exchanger 11 filters + Cold Shutter	Test module pass the robustness test (120K, >100000 moving) and oscillation test
Cooling	M1 100K Other telescope and instrument structure: 80-100K Detector 40-50K Passive radiative cooling	No mechanical cooler

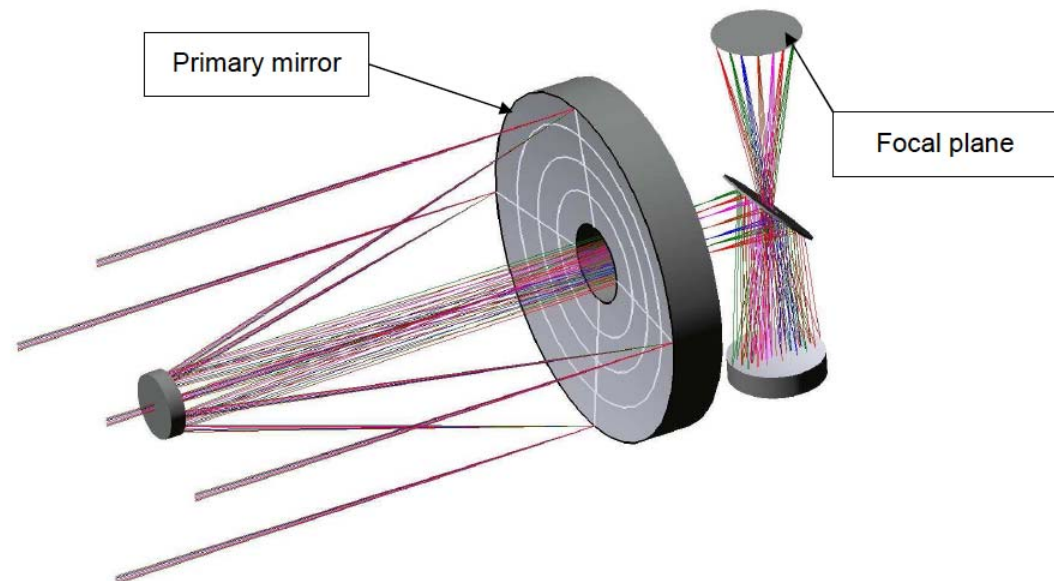
BUS		
SAP	Fixed, 2 wings 1.6kW (BOL) 1.4kW (EOL)	
Sun shield	Al panels + MLI 30/30層	SPICA R&D heritage
Positional Control	STT-IRU Strap down system RW high accuracy (torque-balance)	"internal accuracy" (w/o thermal flexure) <0.03" r.m.s. (300sec) Guiding using the science arrays
Thruster	一液触媒式ブローダウン方式 fuel >162kg thruster 3Nx8 23Nx4	Φ504mm tank 4個
Data link Data processing	X-band (Data, 16Mbps), S-band (telemetry) BUS < 50Mbps (Space Wire 使用) DR 48GB x 2	Ka-band (32Mbps): option

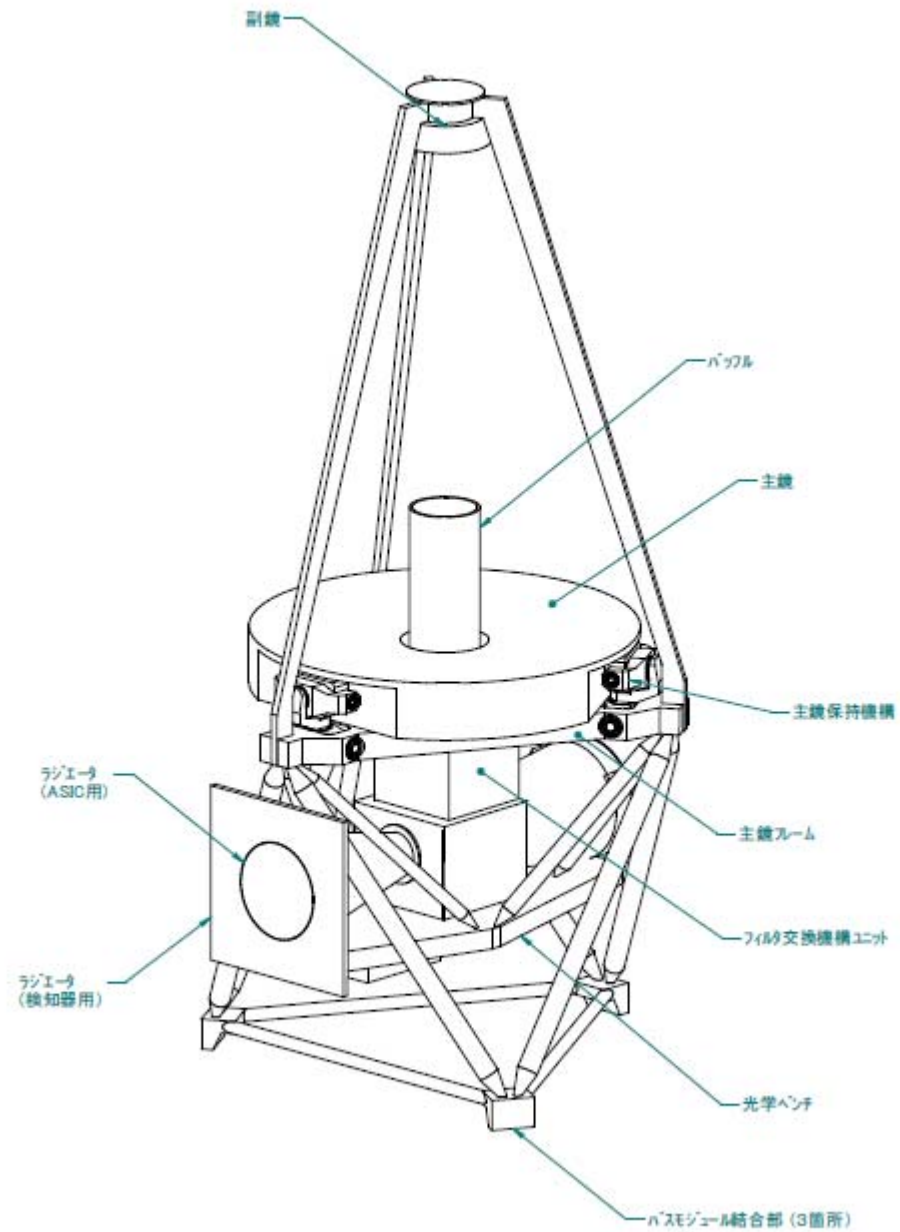
WISH Development: Optical Layout



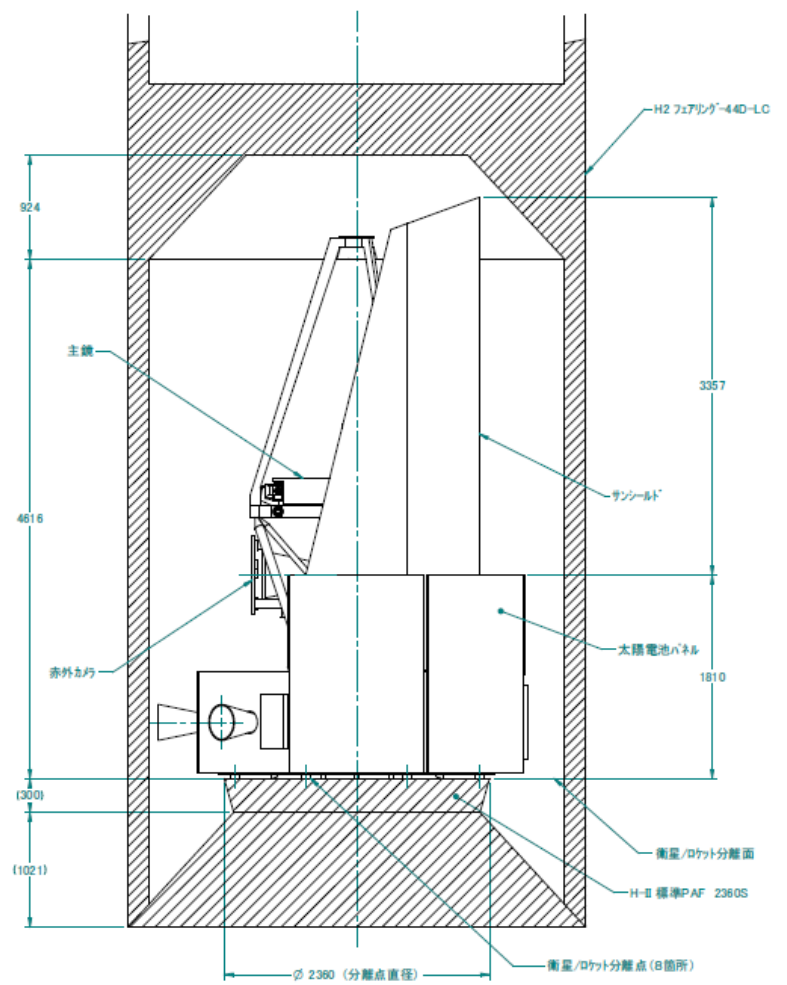
- Very flat focal plane
- Diffraction-limited images to $\phi \sim 50'$ at $1-5 \mu\text{m}$

Yuji Ikeda et al.
(photocoding)



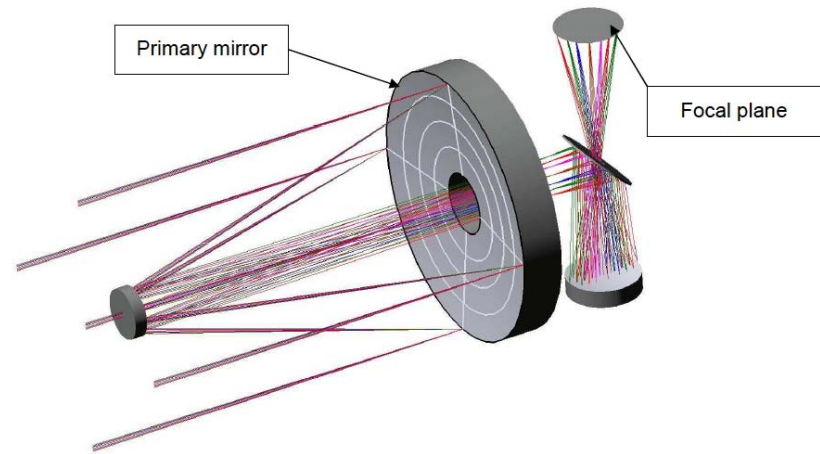
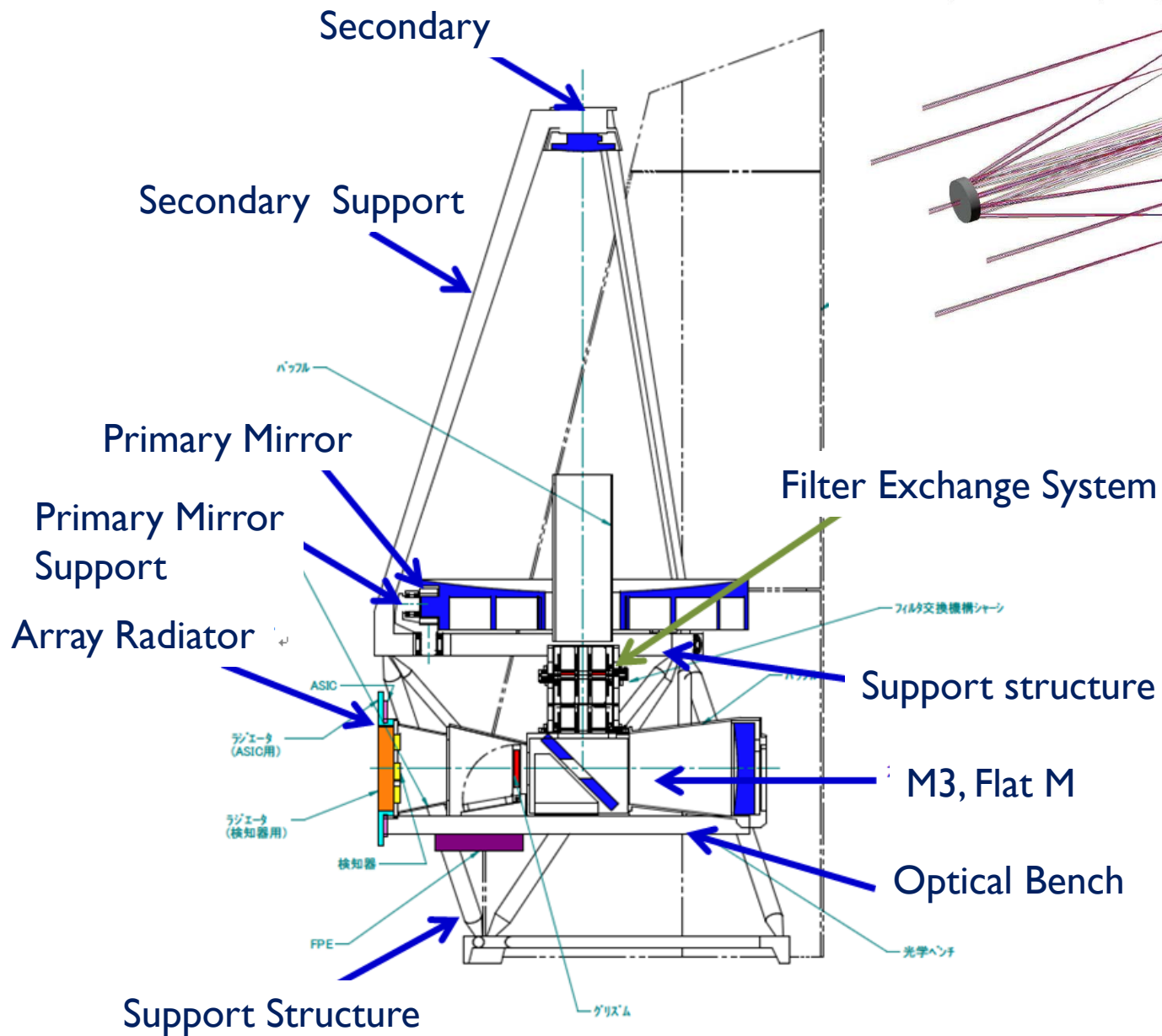


望遠鏡概要

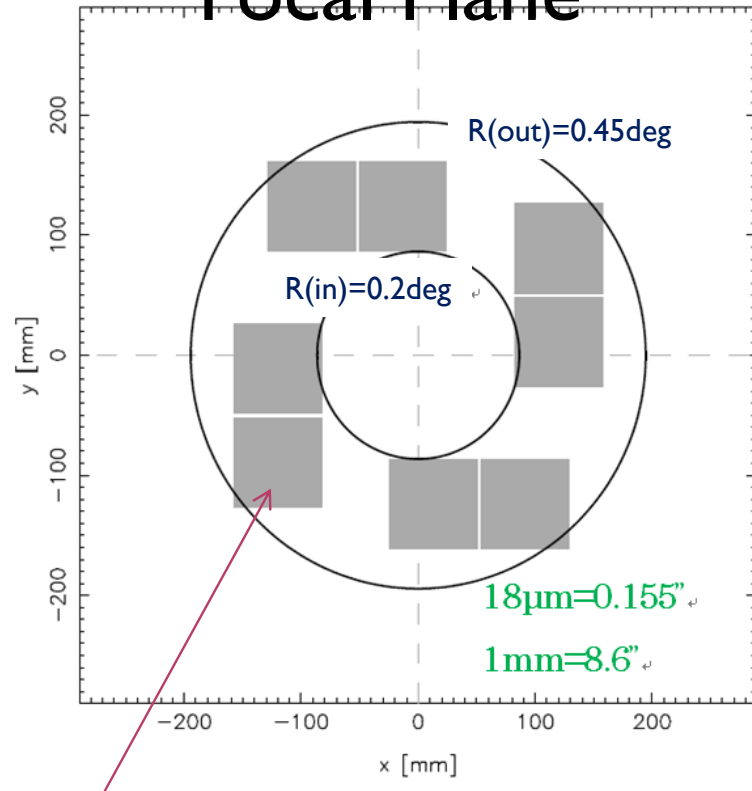


打上時コンフィギュレーション

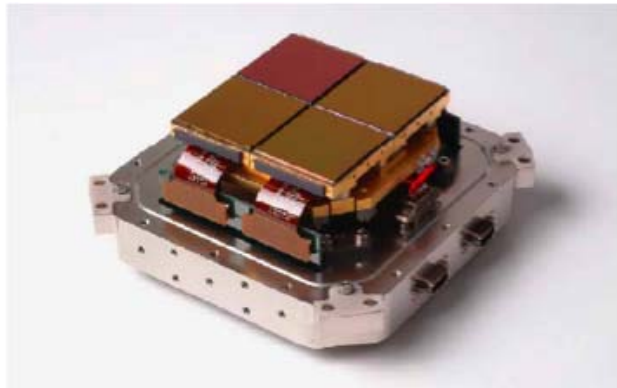
Launch Configuration



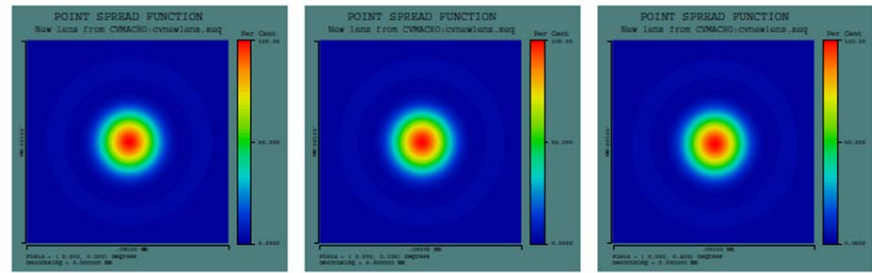
Focal Plane



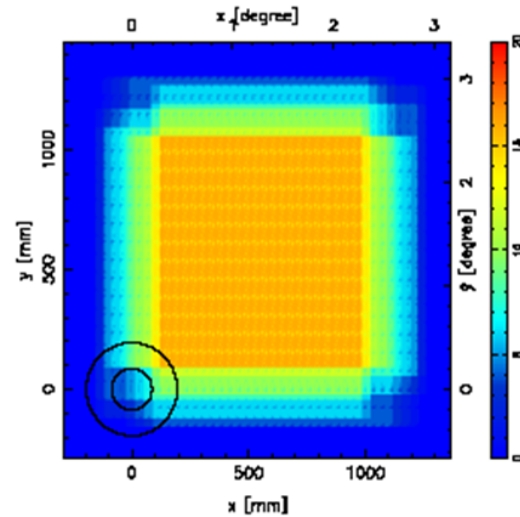
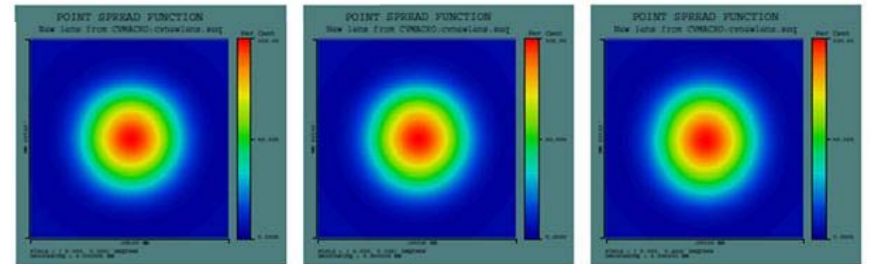
4 x 2kx2k FPA



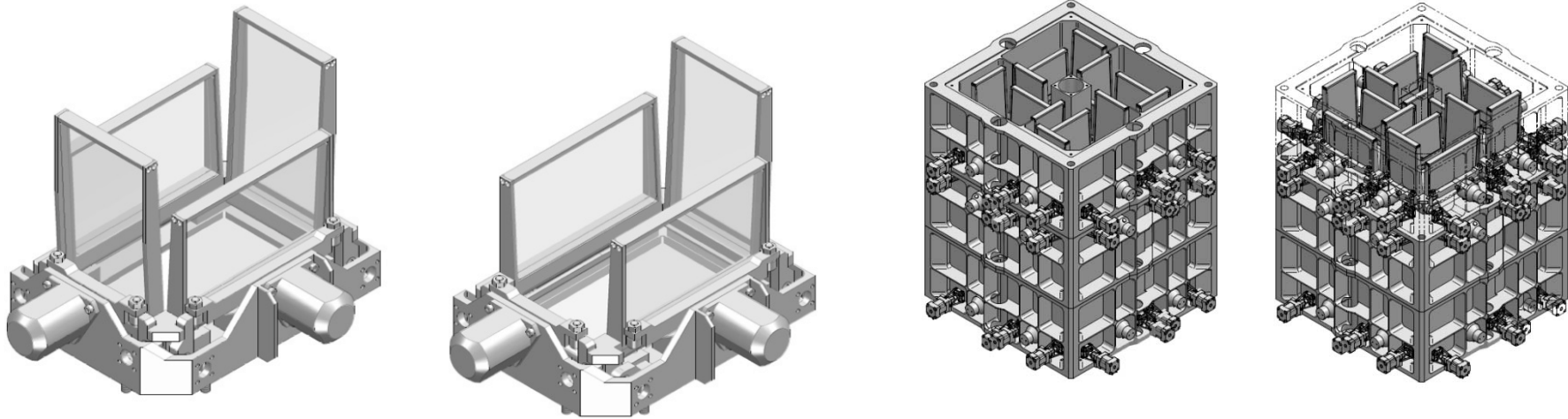
$R=0.2, 0.325, 0.4 \text{ deg @ } 1.25 \mu\text{m}$



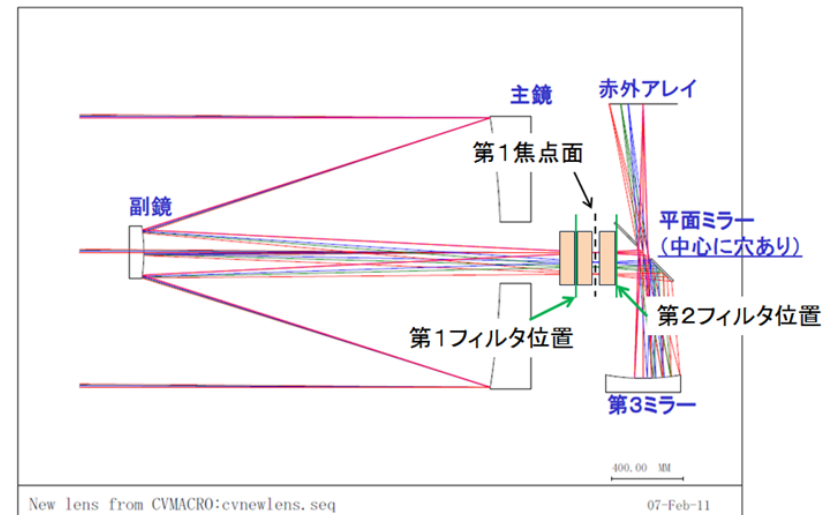
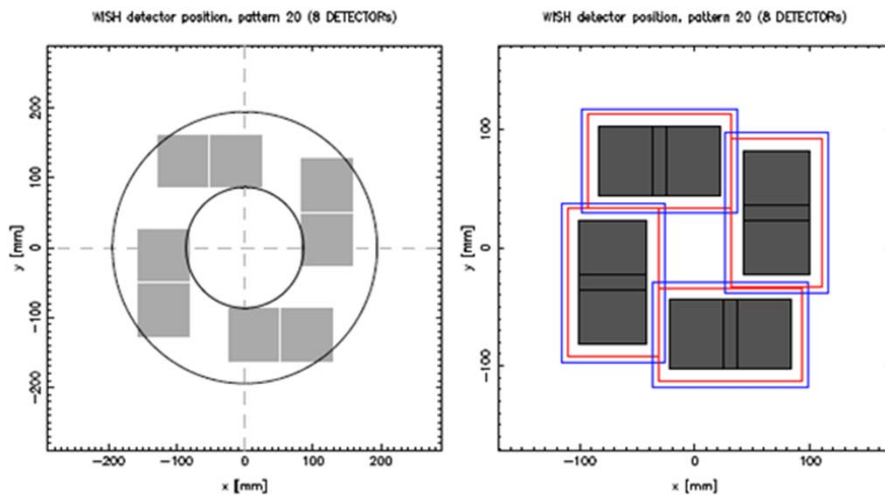
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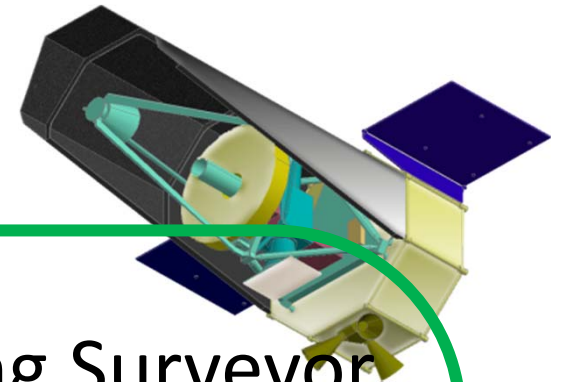
WISH Flip-type Wide-field Filter Exchange System



フィルタ交換機構ASSY 概要図



Summary



- NIR **Deep** and **Wide-field** Imaging Surveyor
- **1.5m** aperture, **0.15''**/pix
- Exploring the 1st generation galaxies
- Dedicated, **~100 deg²**, **28AB (~25nJy)**
- **~10⁴ galaxies at z=8-9, ~3-6x10³ at z=11-12, and ~50-100 galaxies at z=14-17**
- Concept developed under JAXA/ISAS WG to be launched in late 2010's