

A Bayesian algorithm for real-time model selection in caustic-crossing events

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How things work

- Follow-up teams select “promising” MOA/OGLE events, higher cadence or/and better photometry (with e.g. lucky imaging)
- Challenge: new-generation surveys alert ~1700 events/ year
 - How to select best events to follow up?
 - How to model so many events, given the large amount of (human and computational) time required for each event

Microlensing Anomalies

- Following up the “right” anomalies is key
- No simple selection criterion
- Come in a huge variety of shapes and timescales (few hours - few days)
- No way of avoiding computationally intensive modelling

What do we want from modelling

- 2 main purposes:
 - Real-time modelling → predictions of upcoming features, feed information back to telescopes for optimal target selection
 - Post-event modelling: be as thorough and systematic as possible (i.e. explore parameter space)
- Ideally, an algorithm that is good for both
- Several approaches amongst modellers (see talks by V. Bozza, C. Han)

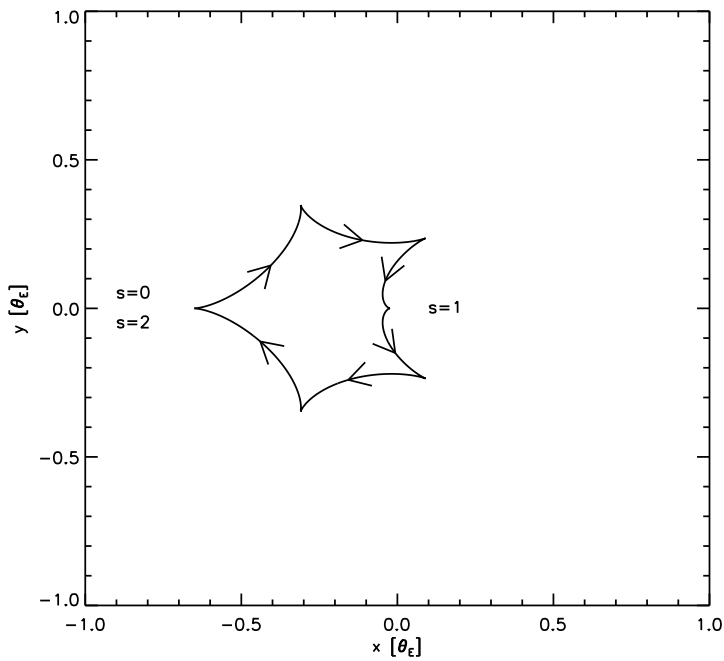
Feature-based parameterisation

- Use parameters that are well constrained by data features
- Caustic-crossing events:
 - Re-parameterised binary-lens lightcurves (Cassan+ 2008, 2010 Kains+ 2009) + MCMC to conduct grid search and generate posterior maps with available data
 - Include as much prior information as possible without biasing modelling

Parameterisation of Cassan (2008)

Locations on the caustic of caustic entry and exit

$$s_{in}, s_{out}$$



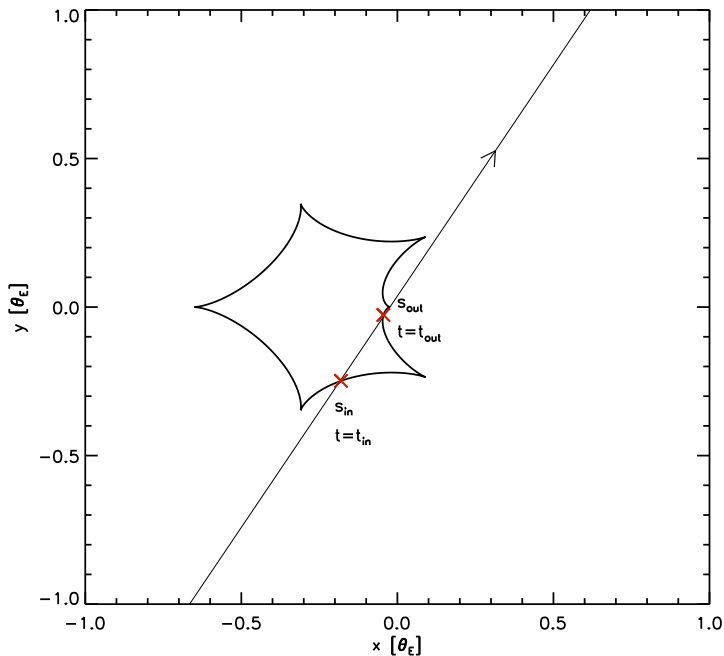
Times of caustic entry and exit

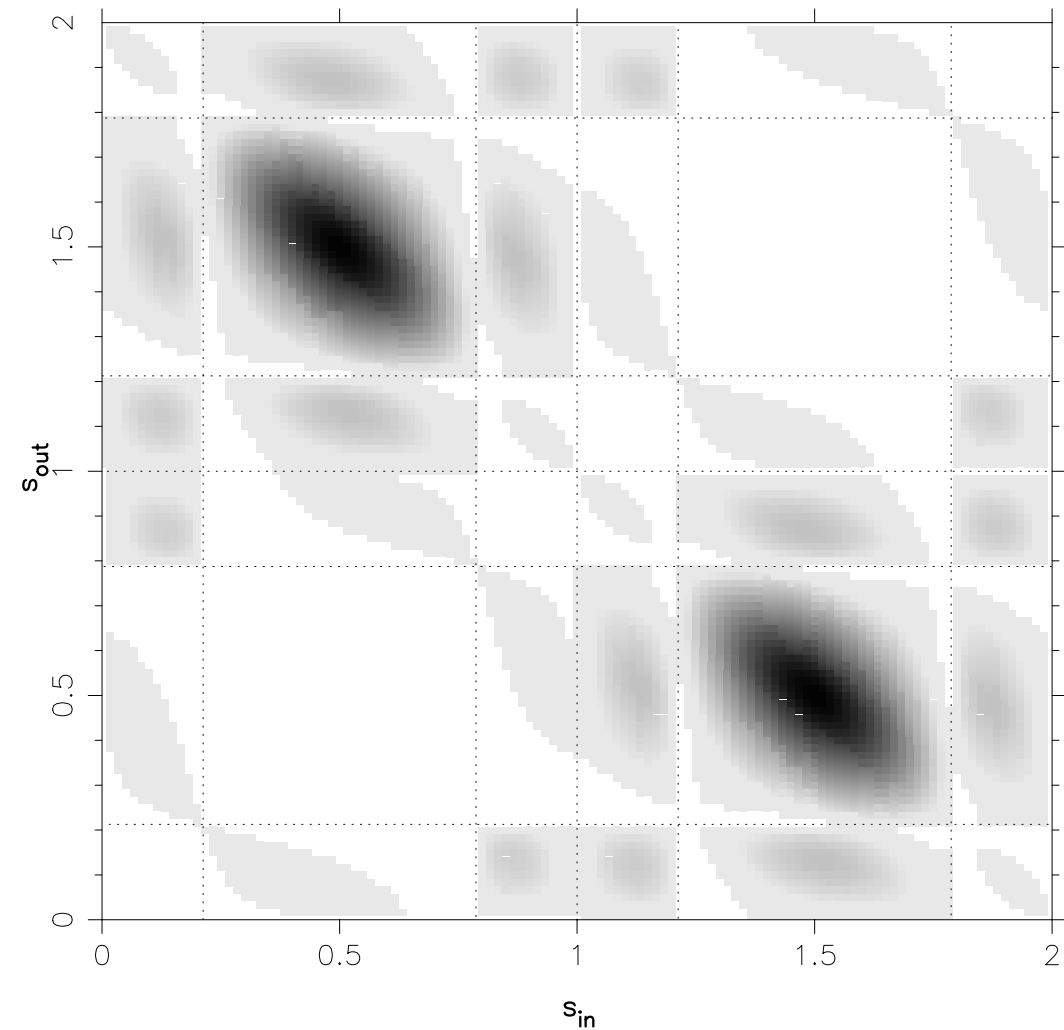
$$t_{in}, t_{out}$$

Duration of crossing Δt_c

$$t_0, t_E, \alpha, u_0, \rho_*(\text{or } t_*), d, q$$

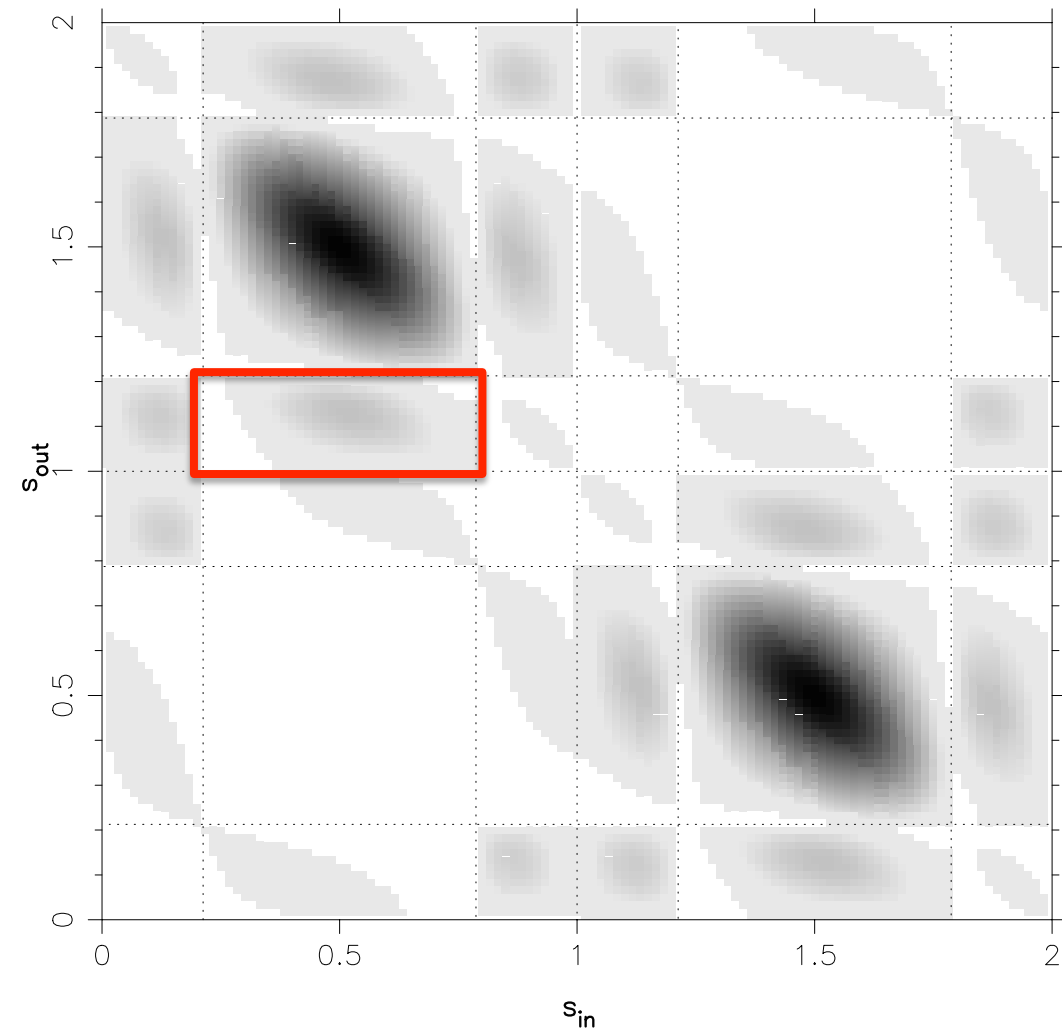
$$\rightarrow t_{in}, t_{out}, s_{in}, s_{out}, \Delta t_c, d, q$$





Assuming uniform distribution of source trajectories and source sizes, this is what a map of a joint prior on (s_{in}, s_{out}) looks like.

We can exploit the structure of these maps to devise a scheme for thorough parameter space exploration

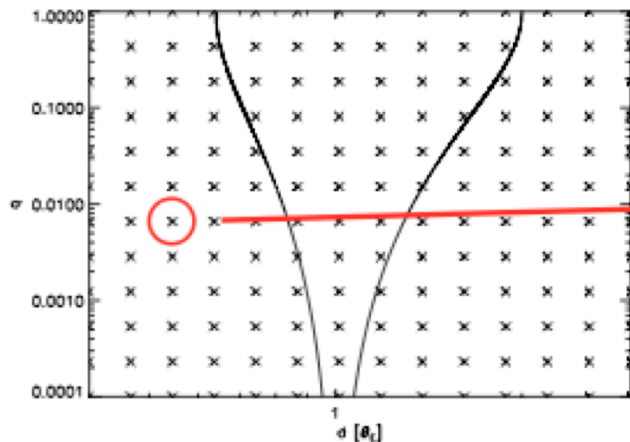


Each 'sub-box' corresponds to the source entering and exiting the caustic on a different caustic fold

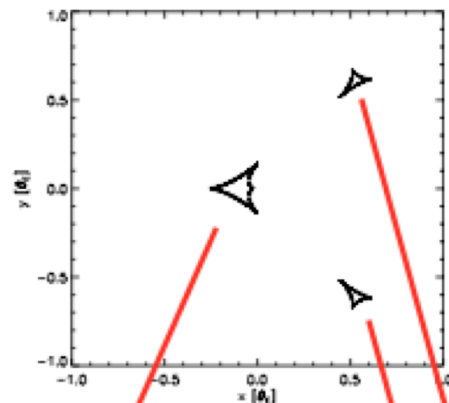
By exploring each sub-box, we can ensure all possible caustic-crossing trajectories are explored for a given caustic

This allows us to conduct a grid search in the (d, q) plane

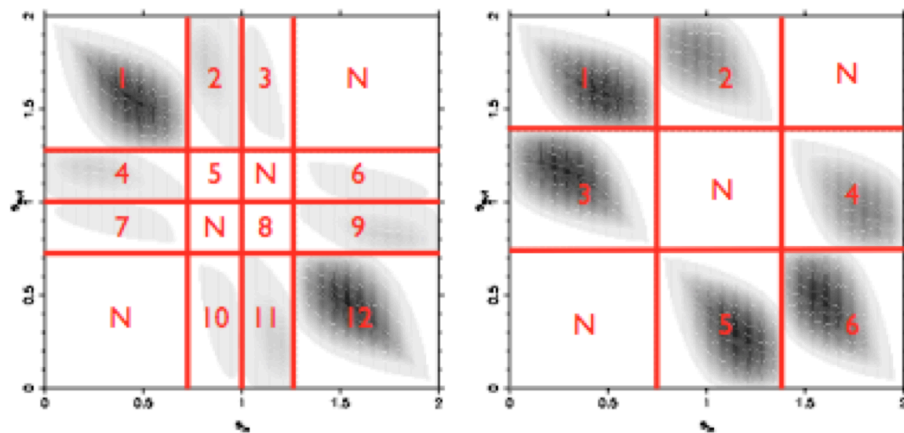
I. (d, q) grid



2. Select caustic

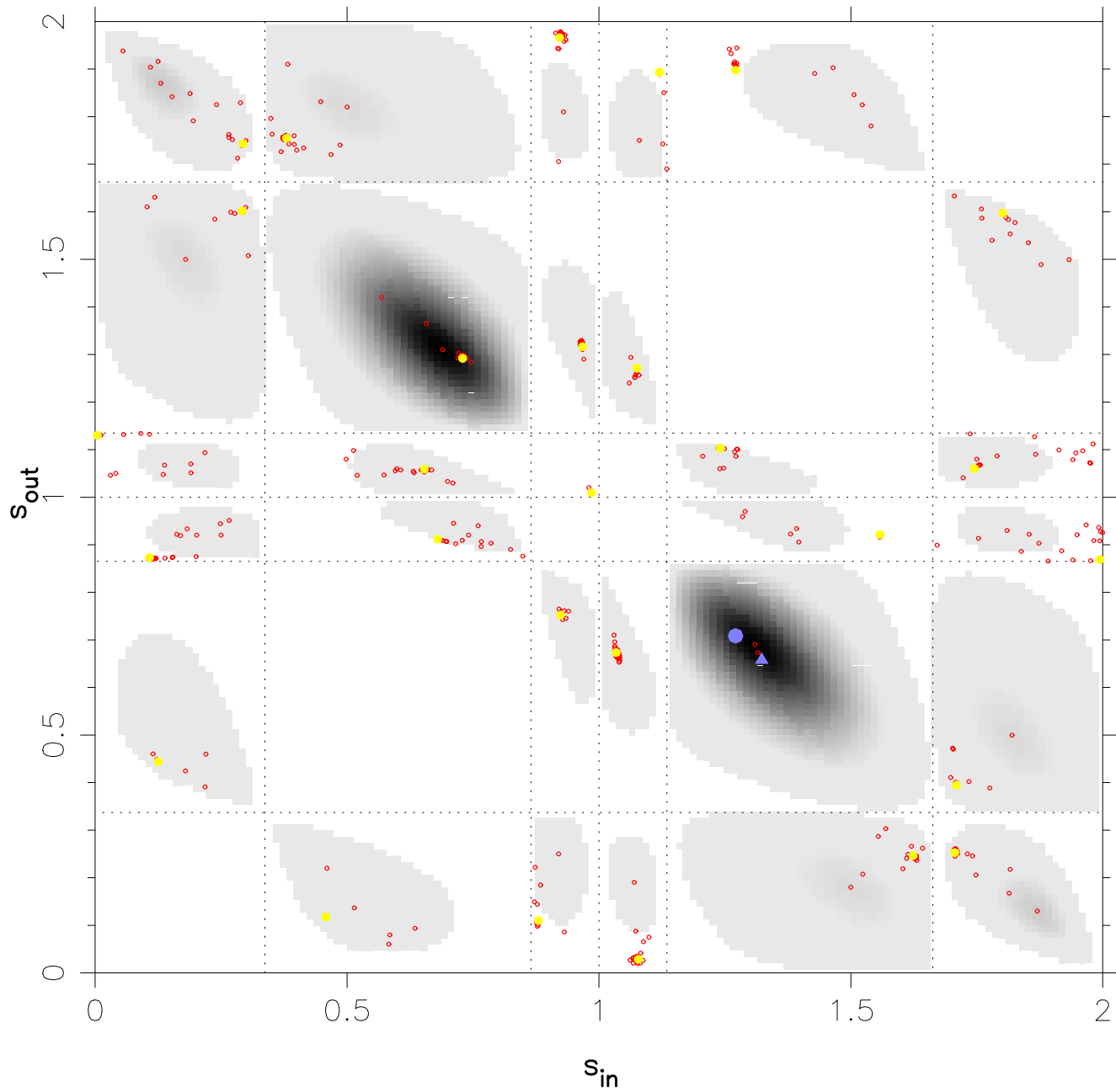


3. Explore $P(s_{in}, s_{out})$ sub-boxes for each caustic

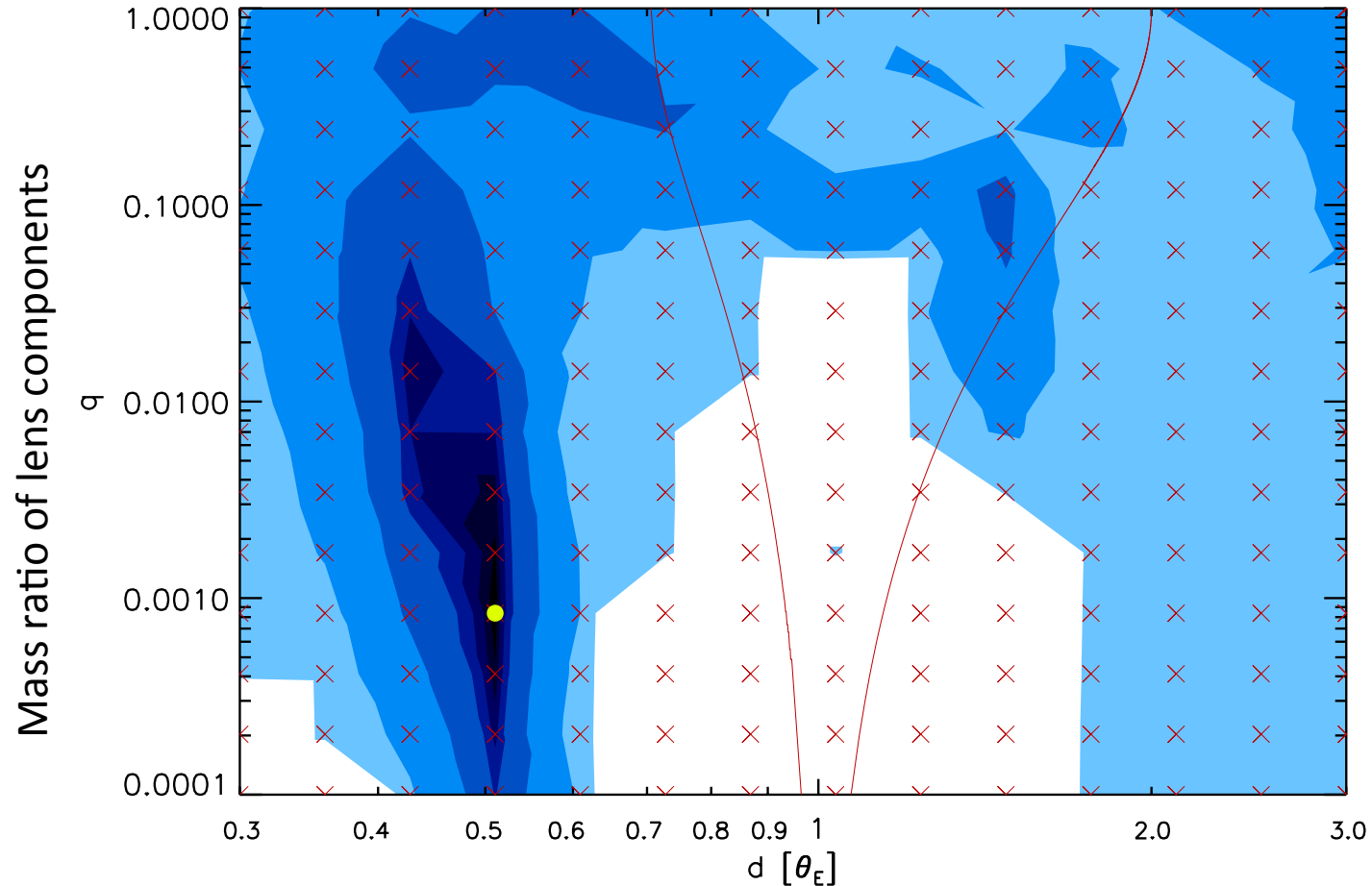


4. Find best fit(s) for each (d, q); iterate





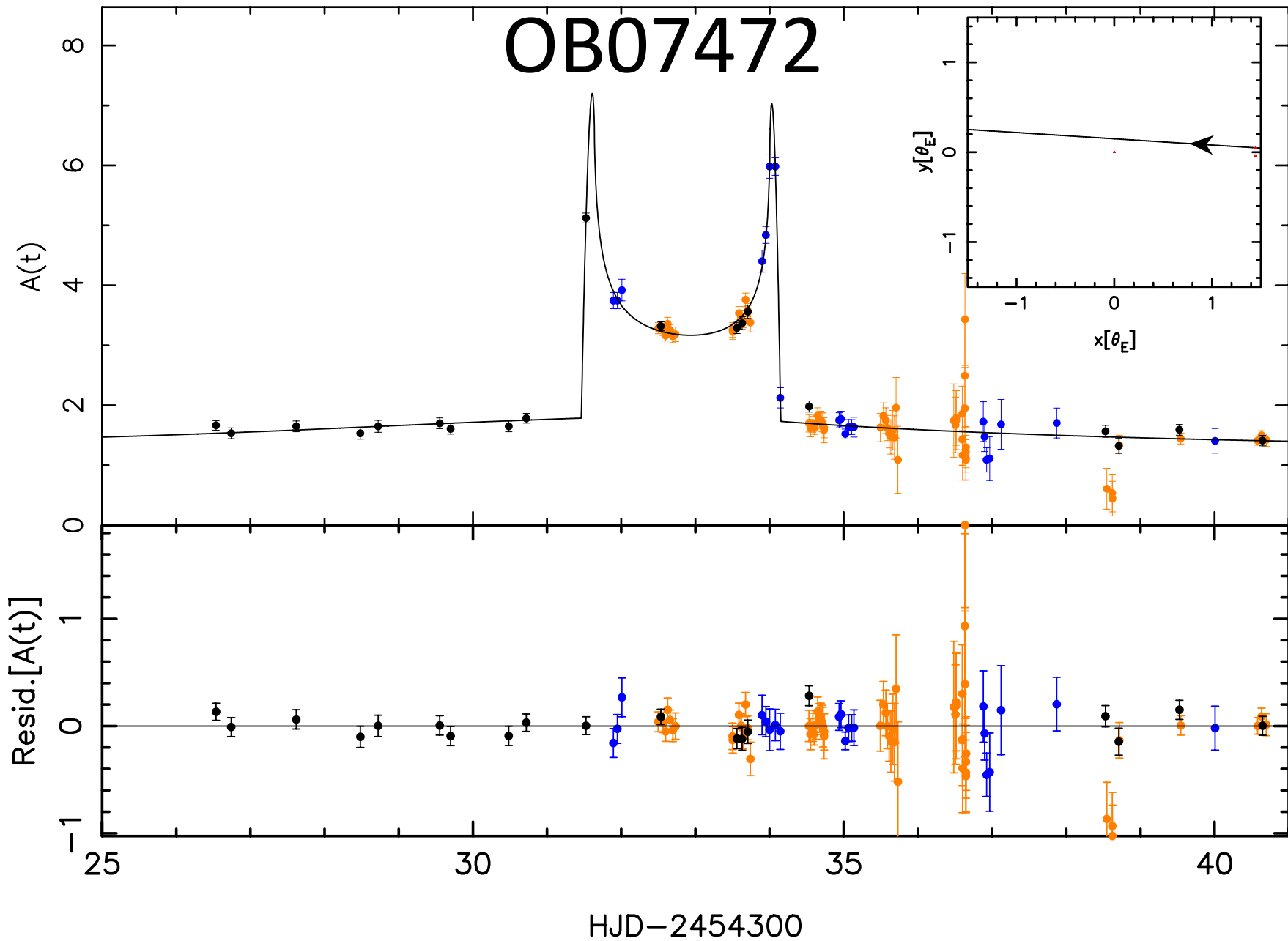
...end up with a “posterior map” of whichever badness-of-fit statistic we choose. E.g. chi2 map:



Separation between lens components (e.g. planetary orbit)

OB070472, first analysed with no priors in Kains+ 2009, also Kains+ in preparation

OB07472

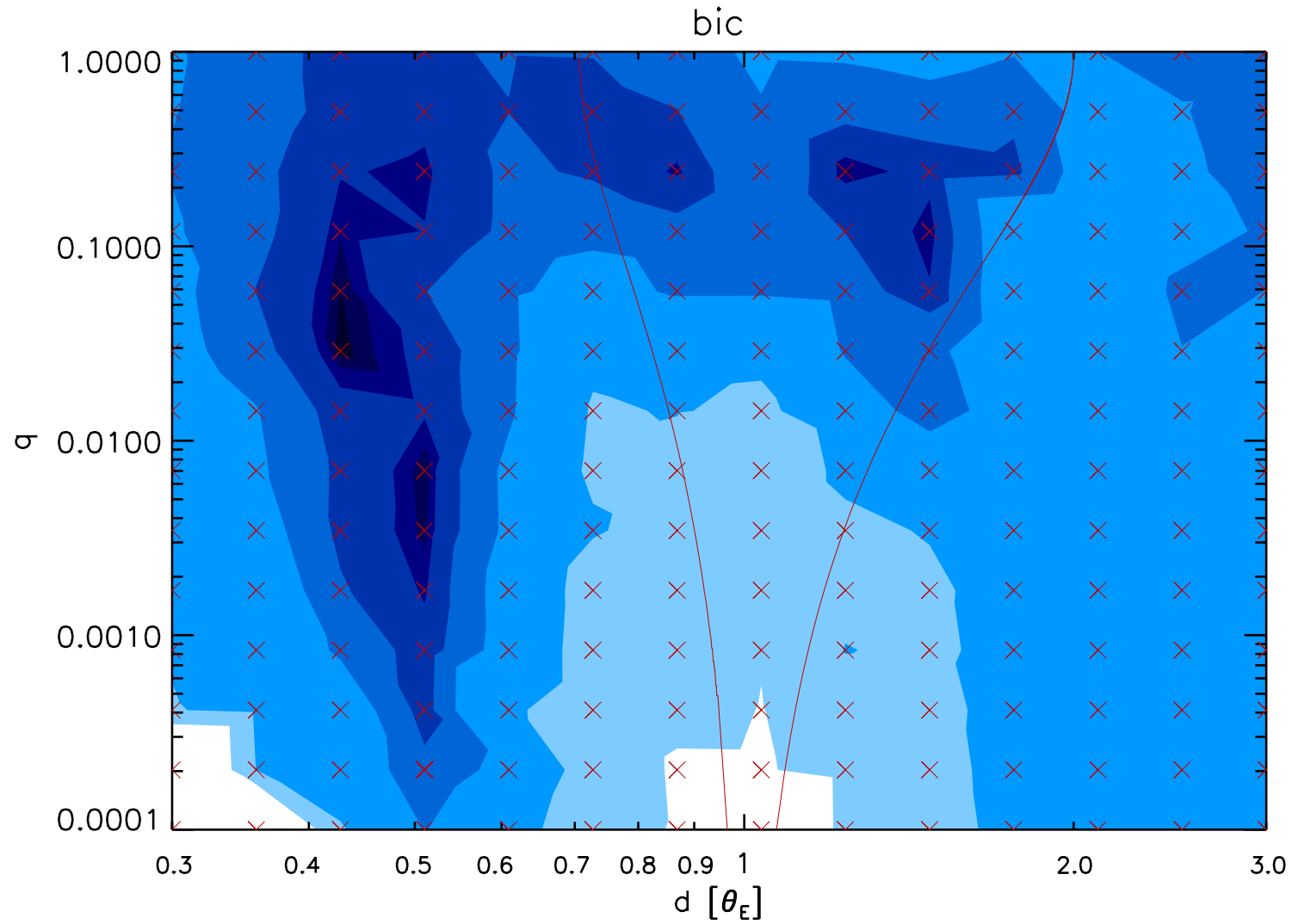


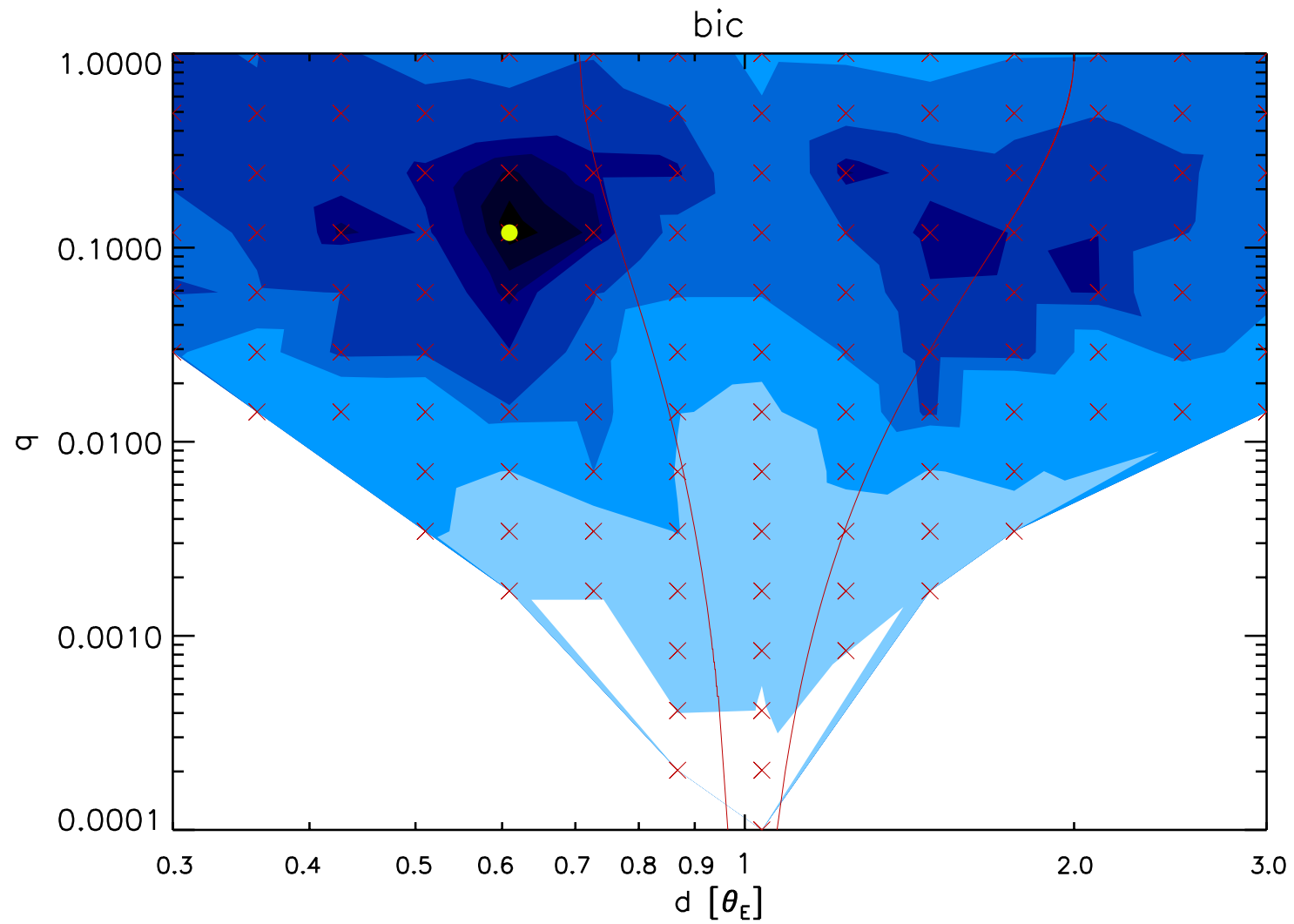
A robust statistic

- Kains+ (2009): best- χ^2 model for OB07472 has extreme t_E (~ 4000 days)
- χ^2 not necessarily the best estimator of true parameters
 - Need for a more reliable badness-of-fit statistic?
 - Include priors
- Some alternatives:
 - Maximum a-posteriori (MAP): Maximise $\chi^2 - 2*\ln(\text{prior})$
 - MAP + volume: take into account parameter space volume
 - Bayesian Information Criterion (BIC): takes into account number of parameters and data points

Choosing priors

- Free to choose prior from any suitable source
- E.g. a prior of timescales could come from a distribution of past observed events or a model distribution (e.g. Wood & Mao 2005)
- We used a joint prior on t_E and ρ_* obtained from a Besançon model simulation (Robin+ 2003)





$$\text{BIC} = X^2 - 2 \ln(\text{Prior}) + N_{\text{eff}} * \ln(N_D)$$

Kains+ in preparation

Conclusions

- Neat way of ensuring the parameter space is explored completely
- Systematic and (nearly) automatic analysis of events
- Improved badness-of-fit statistic allows us to locate more robust best-fit minimum
- Drawbacks: limited to caustic-crossing events for now
- Could extend the parametrisation to include non-crossing caustic approaches etc.